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## GEORGETOWN COLLEGE OBSERVATORY.

# OBSERVATIONS OF VARIABLE STARS 

MADE IN THE YEARS 1884-1890.

## PART I.

THE OBSERVATIONS.
J. G. HAGEN, S. J.,

Director of the Georgetown College Observatory.

ASTRONOMY
cal. or Clawom

## INTRODUCTION.

The following observations were made from 1884 to 1890 by myself and several assistants, first in Prairie Du Chien, Wisconsin, and later in Washington. They were generally confined to the brighter variables, as the instruments used were opera glasses and equatorials of 3 or 5 inches aperture. The earlier observations will show some lack of experience both in their arrangement and in the choice of snitable variables and comparison stars. But the very difficulties encountered in these observations bore ample fruit by showing the necessity of an Atlas of Variable Stars, and by developing and maturing a plan for its construction. As a matter of fact the observations were only discontinued to begin work on the Atlas, in January, 1890, when a donation for a larger instrument was received by the Observatory.

A glance at the observations will show that two methods were employed, viz., the "decimal method" for the first three years, and the "method by steps" for the last three. In the former method the brightness of the variable is estimated in tenths of the interval between two comparison stars; in the latter, usually designated as Argelander's, it is estimated in steps from both comparison stars. The decimal method supposes a well established photometric scale of comparison stars, while the latter rests upon a subjective "step" or unit of light difference. Each method has its own difficulty in the want of constancy of its basis. Experience shows that a large proportion of the stars vary slightly in brightness. Hence any photometric scale must contain on account of this variation what are equivalently accidental errors, however much its constructor may have endeavored to free it from systematic and accidental errors on his own part. On the other hand the "step" is so far from being a constant quantity, that Argelander himself introduced the technical term of the "step value of the evening."

The observations as published show first the original estimates, and then their reduction to a fixed scale either of steps or of photometric magnitudes. Wherever the observations by steps furnished sufficient data, a scale of steps was constructed for all the observations, even those made by the decimal method. This may not be perfectly correct, as the value of the steps may, either in general or for single stars, change in the course of years. Yet it has seemed preferable to make the reductions uniform throughout.

For these reductions of the original extimates to a scale of steps Schönfeld has recommended the use of both the arithmetical and the geometrical proportion. The former is supposed to eliminate erroneous estimates in the actual brightness of the several stars, and the latter removes the difference of the step value of the evening from the assumed mean value. Schöenfeld then advises the taking of the arithmetical mean of the two results thus obtained (Wien, Sitzungsber. vol. 42, p. 154), with a reference to Argelander in Schamacher's Jahrbuch (1844 p. 232). The two proportions may be expressed in algebraic form as follows. Let the observed sequence be : $a \mathrm{~m} \mathrm{R} \mathrm{n} \mathrm{b}$,
$a$ being the brighter and $b$ the fainter comparison star for the variable $R$, and let the steps be counted in the same direction as the magnitudes, viz. increasing from the brighter to the fainter stars, then :

$$
\begin{align*}
& R=\frac{(a+\mathrm{m})+(b-\mathrm{n})}{2}  \tag{I}\\
& R=b-\mathrm{n} \frac{b-a}{\mathrm{n}+\mathrm{m}}=a+\mathrm{m} \frac{b-a}{\mathrm{n}+\mathrm{m}} \tag{II}
\end{align*}
$$

In the latter formula, which uses geometrical proportion, the fraction is what Argelander has called "the step value of the evening."

After these general remarks a few explanations will be required regarding the comparison stars and the observations.

The titles are taken substantially from Chandler's III. Catalogue.
In the tables of comparison stars under the heading "Obs." the letters are given which were used to designate the stars in the observations, and under the next $A S V$. their numbers in the catalogues accompanying the charts of the Atlas. The "Series" in which the variable is contained is mentioned above with the title. When the variable belongs to the IV. Series, the column is left blank for future insertion. $B D$. means the Bonn "Durchmusterung" number. The Sleps were not taken from the Atlas, but derived in the usual way from the observations themselves, except in a few cases which will be specially mentioned. As to the last column, headed "Magn." the following principles have been followed according to the Series of the Atlas to which the variable belongs. For the stars of Series I, II, III the magnitudes are those computed for the Atlas. When a particular comparison star is not found in the Atlas, because lying outside the chart, the BD, magnitude is given in parenthesis. For stars of Series IV, which is not yet published, the BD. magnitudes are given. The catalogues of Series V contain three columns of photometric magnitudes. Here, however, instead of giving the mean of these it was thought best to take the H. P. magnitudes alone.

The table of comparison stars is followed by a few explanatory Notes.
The columns of the Observations require but a few remarks. The four numbers I-IV under Sky are the usual notation for the transparency of the sky, I denoting very good, and IV bad. Disturbing moonlight (not the age of the moon) is denoted in three intensities, by one, two or three signs D.

The passage from the decimal method to that by steps is pointed out in the column Comparisons. For the decimal method, moreover, an abbreviated notation has been used, whose different appearance makes the change of method obvious. Thus the first observation of U Cephei is : $b 3 c$, which is abbreviated from : $b 3 \mathrm{U} 7 c$, and means that the variable U is 0.3 of the interval $c-b$ fainter than $b$, or 0.7 brighter than $c$. The rule of abbreviation is, that the second figure (the complement of the first to 10) and the letter of the variable, are always omitted. Consequently the notations: $d 0 n, a 10 d$ mean: $d 0 \mathrm{U} 10 n$, a $10 \mathrm{U} 0 d$, or that the variable is equal to $d$. Evidently the decimal method coincides with that by steps, whenever the observer estimates 10 steps between the two comparison stars. In the same column the signs ! and ? are not later insertions, but indicate certainty or doubt at the time of observation. The different seasons of observations are separated by horizontal lines, to make the discontinuity of the light curve more apparent to the eye.

The headings $I, I I$ and Mean, refer to the two formulas given above. It is evident, that only the geometrical proportion (formula II) can be applied to the decimal method. Numbers in parenthesis mean that they have been found to bear internal evidence of erroneous or poor observation. They have generally received one-half of the weight of the others. Attention must be called to the use of the signs $>$ and <in the columns headed : Comparisons and Mean. In the former column the notation $\mathrm{T}<c$ for example means that the variable T is fainter than the comparison star c. If the step assigned to $c$ is e. g. 27.8, this observation is expressed in the other column thus: $>27.8$, meaning that the number of steps belonging to T is greater than 27.8 .

The Remarks refer to the original record, whilst critical discussions were entered later as foot notes. That the remarks "seeing poor" or "difficult" are not contradictory to the designation I or II of the sky, is well-known to observers; they usually refer to causes not apparent in the sky.

The Julian Day and the brightness in steps from the preceding column will generally be sufficient to plot the light curves and to compute the phases and periods of the variables, except those of the Algol type. For these the hours and minutes will be needed, and can be taken from the second column.

In the column Remarks the word "Eph." denotes the ephemeris published annually in the V.J.S., and shows that the observations of the Algol stars were generally not arranged with the purpose of obtaining the exact times of the minima.

Finally mention should be made of the fact, that the copying and computing of these observations has been done by six or seven different persons at various times, and that, for this reason, small errors may have escaped notice. Larger errors have probably been detected, as they would cause a break in the sequence of the numbers.

Wasiington, D. C., December 8, 1901.

J. G. HAGEN, S. J.

I. Observations Made by J. G. Hagen, S. J., from 1884 to 1888.

These observations were made in Prairie-Du.Chien, Wisconsin, while teaching in the college there. This will explain why the time of observation was usually confined to the regular hours between 8 and $10 \mathrm{p} . \mathrm{m}$. The instrument was a 3 inch telescope by Merz, of good definition, and the eyepiece employed had a power of 25 diameters, with a field of over one degree. Many of the observations were independently repeated by a student, George Zwack, some 3 or 4 minutes later. These estimates are distinguished by the letter (Z). As an appendix to these observations are added those of Nova Aurigae made later in 1892, at the Georgetown College Observatory.

## U Cephei

Series IV.
(1900) $0^{\mathrm{h}} 53^{\mathrm{m}} 23^{\mathrm{s}}\left(+5^{\mathrm{s}} .09\right) ; \quad+81^{\circ} 20^{\prime} .2\left(+0^{\prime} .33\right)$

Period: $2^{\mathrm{d}} 11^{\mathrm{h}} 49^{\mathrm{m}} \pm ;$ Variation: $7^{\mathrm{m}} 1-9^{\mathrm{m}} 2$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| b |  | $+81^{\circ} 13$ | 0.0 | 6.5 BD. |
| c |  | +81 18 | 6.5 | 7.6 " |
| e |  | +8130 | 10.7 | 8.3 ، |
| a |  | +8127 | 13.2 | 8.6 " |
| d |  | +8122 | 17.4 | 9.2 " |
| n |  | $+8126$ | 23.8 | 9.5 " |

Notes:
This Algol star was not observed systematically with the view of obtaining complete determinations of the Minima, but rather for practice, to get a general knowledge of its variations. The observations may be utilized for the study of the light curve, now that the period is well determined.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:



## METHOD BY STEPS:



| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 87 Oct. 12 | $15{ }^{\text {b }} 52^{\mathrm{m}}$ |  | d 1 U5n | 18.6 | 18.5 | 18.3 | 10557 |  |
|  |  |  | a 4 U 5 n | 18.9 | 17.9 |  |  |  |
| ، | 164 |  | d 2 U 5 n | 19.1 | 19.2 | 19.2 | " |  |
| " | 1613 |  | d 2 U 5 n | 19.1 | 19.2 | 18.8 | " | Eph. ${ }^{\frac{1}{2} \mathrm{~h}}$ before Min. |
| Nov. 11 | 154 | I | d2 2 U 4 n | 19.6 | 19.4 | 19.5 | 587 | ، $\quad \frac{1}{2} \mathrm{~h}$ after Min. |
| " | 1518 | I | d 2 U 5 n | 19.1 | 19.2 | 18.8 | 6 | , ${ }_{2}$ after Min. |
|  |  |  | a 5 U 5 n | 18.5 | 18.5 |  |  |  |
| " | 1534 | " | a 4 U 0.5 d | 17.1 | 17.0 | 17.1 | " |  |
| " | 168 | " | a 3 U 2 d | 15.8 | 15.7 | 15.8 | " |  |

782 R Arietis Series II.
(1900) $2^{\mathrm{h}} 10^{\mathrm{m}} 25^{\mathrm{s}}(+3.40) ;+24^{\mathrm{s}} \quad 35.5 \quad(+0.28)$

Period : 186.55; Variation: $8^{\frac{\mathrm{m}}{\frac{1}{2}}}-12^{\frac{\mathrm{mp}}{2}}$.
Comparison Stars :

| Obs. | ASV. | BD . | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| a | 1 | +24 ${ }^{\circ} 329$ | 0.0 | [6.0] BD. |
| k | 2 | +23303 | 7.2 | [6.5] BD. |
| g | 4 | +23306 | 15.0 | 8.9 |
| b | 7 | +24327 | 18.0 | 9.4 |
| $f$ | 5 | +24,334 | 20.8 | 9.2 |
| e | 6 | +24333 | 22.2 | 9.4 |
| c | 9 | +24331 | 25.8 | 9.6 |
| h | 14 | - ........ | 30.1 | 10.0 |
| d | 13 | $\cdots$ | 31.5 | 10.0 |

Notes :
The last column indicates that comparison star $b$ was estimated brighter with the 3 -inch telescope than with larger instruments in later years. The sequence in the Atlas is as follows: $\mathrm{g}, \mathrm{f}, \mathrm{e}, \mathrm{b}, \mathrm{c}$. Whatever the cause of the discrepancy may be, it was thought more correct to reduce the observations by the above scale, which is derived from the observations themselves.

The comparison star $g$ was in later years suspected of slight variations.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |

## DECIMAL METHOD:



*) The two observations are contradictory.
**) b 1 h can have little weight, since the interval $\mathrm{h}-\mathrm{b}=12.1$ steps is too large.

| $1800+$ |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 86 Jan. | 6 | 15.3 | III | f 4 e |  | 21.4 | 22.0 | 09913 |  |
|  | 30 | 15.7 | II-III | f 3 e |  | 22.5 |  | 937 |  |
|  |  |  |  | a 6 c |  | (15.5) | 13.1 |  |  |
|  | 2 | 15.6 | III | a 7 f |  | 14.6 |  |  |  |
| Feb. |  |  |  | a 7 c |  | (18.1) | 14.0 | 940 |  |
|  |  |  |  | a 8 g |  | 12.0 |  |  |  |
|  | 7 | 15.6 | I | k 6 g |  | 11.9 | 12.0 | 945 |  |
|  |  | 15.7 | I | a 8 g |  | 12.0 |  |  |  |
|  | 22 |  |  | k 6 g |  | 11.9 | 12.0 | 960 |  |
|  | 25 | 15.7 | I | a 8 g k 8 g |  | 12.0 13.4 | 14.8 | 963 |  |
|  |  |  |  | k 78 |  | (17.7) |  |  |  |
| Mar. |  | 15.7 | I | k 8 g |  | 13.4 | 13.4 | $\begin{aligned} & 968 \\ & 971 \end{aligned}$ |  |
|  | 5 | 15.1 | I | k 9 g |  | 14.2 | 15.9 |  |  |
|  | 25 | 14.2 | I |  |  | (19.2) | 27.3 | 991 |  |
|  |  |  |  | c 6 d |  | 29.2 | 27.3 | 991 |  |

## METHOD BY STEPS:



*     * *) The journal has R1g3f, which would at that time have been to the observer a very unusual way of recording.

| $1800+$ | Gr. M. 'T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 88 Jan. | 16.4 | I | c 1 R 3 h ! | 27.0 | 26.9 | 27.0 | 10645 | Near horizon |
|  | 14.8 | I | f 3 R 2 c ! | 23.8 | 23.8 | 23.8 | 648 |  |
|  | 14.7 | I D | f 2 R 3 c | 22.8 | 22.8 | 22.8 | 654 |  |
| Feb. | 15.1 | I | a 7 R 3 g | 9.5 | 10.5 | 10.8 | 675 |  |
|  |  |  | a 7 R 5 f | 11.4 | 12.1 |  |  |  |
|  | 14.1 | I | a 7 R 4 m | 9.0 10.0 | 9.5 | 9.6 | 682 |  |
| Mar. $\begin{array}{r}6 \\ 13\end{array}$ | 14.8 | I | k 4 R 3 g | 11.6 | 11.7 | 11.7 | 703 |  |
|  | 14.7 | I | k 5 R 3 g | 12.1 | 12.1 | 12.1 | 710 |  |
| Aug. 12 | 16.1 | I | k 5 R 3 g | 12.1 | 12.1 | 12.1 | 862 | " ${ }^{\text {a }}$ |
| Sept. 6 | 15.3 | I | R 4 g | 11.0 |  | 11.0 | 887 | ، ، |
| Oct. 7 | 16.8 | I | b $2 \underset{\text { R }}{\text { R }}$ f 3 e | 19.6 20.8 | 19.9 | 20.1 | 918 |  |

8 I 4 S Persei Series III.

$$
(1900) \quad 2^{\mathrm{h}} 15^{\mathrm{m}} 41^{8}(+4.27) ; \quad+58^{\circ} 7^{\prime} .8\left(+0 .^{\prime} 28\right)
$$

Long period; Variation: $8_{\frac{1}{2}}^{\frac{\mathrm{L}}{2}}-12^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | B D. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  | 1 | $+58^{\circ} 471$ | 0.0 | 7.8 |
| f | 4 | +58467 | 3.3 | 8.2 |
| e | 5 | $+58452$ | 4.3 | 8.3 |
| k | 7 | +57549 | 5.3 | 8.5 |
| d | 9 | +58457 | 8.7 | 8.8 |
| b | 18 | +57557 | 14.5 | 9.5 |
| c | 36 | $\cdots$ | 19.1 | 10.9 |
| a | 38 | ..... .. | 23.1 | 11.0 |

Notes:
The records "invisible" are very important in the case of this star, as without them it would have been very difficult to prove that the period in Chandler's Cat. I. was too short.

These observations were continued in Washington for a short while. See below No. III.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:


*) The Journal has d 6 S with a correction into e 6 S ; which must have been made soon after.

(1900) $2^{\mathrm{h}} 42^{\mathrm{m}} 45^{\mathrm{s}}(+3.34) ;+17^{\circ} 5^{\prime} .5 \quad\left(+0^{\prime} .25\right)$

Period: $313^{\text {d }}$ (periodic inequal.); Variation : $8 \frac{1}{2}^{n}-9 \frac{1}{2}^{x}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magı. |
| :---: | :---: | :---: | :---: | :---: |
| h |  | $+16^{\circ} 342$ | 0.0 | 7.8 BD. |
| g |  | +16346 | 6.7 | 8.7 " |
| e |  | +16345 | 8.7 | 8.6 " |
| c |  | +16348 | 12.7 | $8.8$ |
| d |  | +17440 | 14.7 | 8.9 " |
| a |  | +16350 | 18.9 | 9.5 " |
| b |  | +16347 | 22.2 | 9.5 " |

## Notes:

The comparison star $g$ was suspected of variability, as it appeared at times decidedly fainter than $e$. The changes seem to depend on the season of the year, and consequently on the position of the observer.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:

| 83 Sept. | 26 | 16.6 |  | e 3 d | 10.5 | 10.5 | 09080 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 30 | 15.8 | I | e 4 d | 11.1 | 11.1 | 084 |
| Oct. | 6 | 15.4 | I | e 8 d | 13.5 | 13.5 | 090 |
| Nov. | 1 | 15.4 | I | e 8 d | 13.5 | 13.5 | 116 |
|  | 3 | 15.0 | I | d 4 a ! | 16.4 | 16.4 | 118 |
|  | 6 | 15.1 | II D | e 8 d ! | 13.5 | 13.5 | 121 |
|  | 7 | 14.8 | I D | d 1 a | 15.1 | 15.1 | 122 |
|  | 18 | 14.9 | II | d 2 a | 15.5 | 15.4 | 133 |
|  |  |  |  | d 1.5 a (Z) | 15.3 |  |  |
|  | 26 | 15.5 | I | d 4 a | 16.4 | 16.8 | 141 |
|  |  |  |  | d 6 a (Z) | 17.2 |  |  |
|  | 27 | 15.4 | III | d 7 a | 17.6 | 17.4 | 142 |
|  | 29 | 14.5 | II | d 6 a ${ }^{\text {d } 7 \text { a }}$ | 17.2 | 17.3 | 144 |
|  | 29 |  | II | d 5 a (Z) | 16.9 | 17.0 | 144 |
| Dec. | 21 | 14.8 | III | d 4 a | 16.4 | 16.8 | 166 |
|  |  |  |  | d 6 a (Z) | 17.2 |  |  |
|  | 25 | 16.1 | III | d 4 a | 16.4 | 16.4 | 170 |
|  |  |  |  | d 4 a (Z) | 16.4 |  |  |
|  | 28 | 14.8 | II | d 4 a ${ }^{\text {d }}$ a $(Z)$ | 16.4 16.9 | 16.5 | 173 |
| 84 Jan. | 3 | 14.9 | III | d 1 a | 15.1 | 14.9 | 179 |
|  |  |  |  | d 0 a (Z) | 14.7 |  |  |
|  | 5 | 15.0 | I D | e 8 d | 13.5 | 12.9 | 181 |
|  |  |  |  | e 6 d (Z) | 12.3 |  |  |
|  | 15 | 14.9 | I | e 5 d | 11.7 |  | 191 |
|  | 17 | 14.5 | I | e 4 d ${ }_{\text {d }}$ | 11.1 | 11.7 | 193 |
|  | 19 | 14.2 | II | e 8 d ( | 13.5 | 12.6 | 195 |
|  |  |  |  | e 5 d (Z) | 11.7 |  |  |
|  | 20 | 15.1 | II | e 3 d | 10.5 | 10.8 | 196 |
|  |  |  |  | e 4 d (Z) | 11.1 |  |  |
|  | 21 | 14.5 | II |  | 12.9 11.1 | 12.0 | 197 |
|  | 23 | 15.3 | I | e 54 d | 11.7 | 11.4 | 199 |
|  |  |  |  | e 4 d (Z) | 11.1 |  |  |
|  | 24 | 14.4 | II | e 3 d | 10.5 | 11.1 | 200 |
|  |  |  |  | e 5 d (Z) | 11.7 |  |  |
|  | 30 | 15.1 | II | e 7.5 d | 13.2 | 13.2 | 206 |
|  | 31 | 14.4 | III | e 8 d | 13.5 | 12.3 | 207 |
| Feb. | 14 | 14.7 | III | e 4 d d ( 2 ) | 11.1 13.5 | 12.6 | 221 |
|  |  |  |  | e 5 d (Z) | 11.7 |  |  |
|  | 19 | 14.9 | I | $\begin{aligned} & \text { e } 10 \text { d } \\ & \text { e } 8 d(Z) \end{aligned}$ | 14.7 13.5 | 14.1 | 226 |
| Sept. | 17 | 16.3 | II-III | e 9 c | 12.3 | 13.2 | 437 |
|  |  |  |  | e 8 d | 13.5 |  |  |
|  |  |  |  | e 5 a | 13.8 |  |  |
|  | 24 | 16.4 | I | e 8 d e 6 a | 13.5 | 14.2 | 444 |
|  |  |  |  | e 6 a | 14.8 |  |  |

\begin{tabular}{|c|c|c|c|c|c|c|c|c|c|}
\hline 1800 \& \& Gr. M. T. \& Sky \& Comparisons \& I \& 11 \& Mean \& $2400000+$ \& Remarks <br>
\hline \multirow[t]{2}{*}{84 Oct.} \& 12 \& 15.5 \& I \& e 10 c
e 5 d
e 3 a \& \& 12.7
11.7
11.7 \& 12.1 \& 09462 \& <br>
\hline \& 17 \& 15.1 \& I \& e 10 c \& \& 12.7 \& 11.6 \& 467 \& <br>
\hline \multirow[t]{5}{*}{Nov.} \& 8 \& 14.9 \& II \& e 3 d \& \& 10.5
13.5 \& 14.0 \& 489 \& <br>
\hline \& \& \& \& c 3 a \& \& 14.5 \& \& \& <br>
\hline \& \multirow[t]{2}{*}{18
20} \& 15.0 \& I \& d 4 a \& \& 16.4 \& 16.4 \& 499 \& <br>
\hline \& \& 15.7 \& I \& c 6 ca \& \& 16.4 \& 14.8 \& 501 \& <br>
\hline \& \& \& \& c 4 a \& \& 15.2 \& \& \& <br>
\hline Dec. \& 22 \& 14.9 \& II \& c 7 a \& \& 17.0 \& 17.5 \& 533 \& <br>
\hline \multirow[t]{2}{*}{85 Jan.} \& \multirow[t]{2}{*}{9} \& \multirow[t]{2}{*}{15.8} \& \multirow[t]{2}{*}{I} \& e 9 a
d 6 a \& \& 17.9
17.2 \& 17.5 \& 551 \& <br>
\hline \& \& \& \& d 4 b \& \& 17.7 \& \& \& <br>
\hline \multirow[t]{6}{*}{Oct.} \& \& 15.6 \& II \& d 3 a \& \& 16.0 \& 16.0 \& 816 \& <br>
\hline \& 8 \& 15.6 \& I \& e 9 d ? \& \& 14.1 \& (14.1) \& 823 \& <br>
\hline \& 9 \& 15.8 \& I \& e 9 d ! \& \& 14.1 \& 14.1 \& 824 \& <br>
\hline \& 13 \& 15.7 \& I \& e 9 d
e 6 a \& \& 14.1
14.8 \& 14.5 \& 828 \& <br>
\hline \& \multirow[t]{2}{*}{15} \& 15.4 \& D \& d 3 a ! \& \& 16.0 \& 16.0 \& 830 \& Difficult <br>
\hline \& \& 15.4 \& \& d 3 a \& \& 16.0 \& 15.9 \& 844 \& *) <br>
\hline \multirow[t]{5}{*}{Nov.} \& 4 \& 15.8 \& II \& e ${ }^{\text {d }} 4 \mathrm{a}$ \& \& 16.4 \& 16.1 \& 850 \& <br>
\hline \& \& \& \& c 5 a \& \& 15.8 \& \& \& <br>
\hline \& 9 \& 15.3 \& I \& c 9 d ! \& \& 14.5 \& 15.5 \& 855 \& <br>
\hline \& 12 \& 15.2 \& II \& d 4 a? \& \& 16.4 \& (16.1) \& 858 \& <br>
\hline \& \& \& \& e 5 a? \& \& 15.8 \& \& \& <br>
\hline \multirow[t]{4}{*}{Dec.} \& 2 \& 15.5 \& I \& d 3 a \& \& 16.0 \& 15.6 \& 878 \& <br>
\hline \& \multirow[t]{2}{*}{6} \& 15.1 \& I \& c 4 a
d 4 a \& \& 15.2
16.4 \& 16.2 \& 882 \& <br>
\hline \& \& \& \& d 2 b \& \& 16.2 \& \& \& <br>
\hline \& \& 15.7 \& I \& c 5 a
d 3 a

a \& \& 16.1 \& 16.0 \& 887 \& <br>
\hline \multirow[t]{4}{*}{86 Jan.} \& 5 \& 15.3 \& I \& e 9 c \& \& 12.3 \& 13.0 \& 912 \& <br>
\hline \& \& \& \& e 7 d \& \& 12.9 \& \& \& <br>
\hline \& \multirow[t]{2}{*}{11} \& 15.4 \& I \& e 5 a
c 1 d \& \& 13.8 \& \& \& <br>
\hline \& \& \& \& c 1 d
c 2 a \& \& 12.9
13.9 \& 13.4 \& 918 \& <br>
\hline \multirow[t]{2}{*}{Jan.} \& \multirow[t]{2}{*}{30} \& \multirow[t]{2}{*}{15.8} \& \multirow[t]{2}{*}{II-III} \& g 0 e \& \& 6.7 \& 7.4 \& 937 \& <br>
\hline \& \& \& \& g 2 d
g 1 c
c \& \& 8.3
7.3 \& \& \& <br>
\hline \multirow[t]{3}{*}{Feb.} \& \multirow[t]{3}{*}{1.} \& 15.4 \& \& $\bigcirc 9 \mathrm{~g}$ \& \& 6.0 \& 6.0 \& 939 \& <br>
\hline \& \& 15.3 \& I \& h 9 g \& \& 6.0 \& 7.3 \& 940 \& <br>
\hline \& \& \& \& h 8 e
h 7 c \& \& 7.0
8.9 \& \& \& <br>
\hline
\end{tabular}

*) The Journal has another estimate: d 7 c , for which no explanation can be found, as it is contradictory to the scale and to the other observations.

| 1800+ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 86 Feb. | 15.0 | I | h 9 g |  | 6.0 | 6.5 | 09941 | $\mathrm{T}>\mathrm{g}!$ |
|  | 15.2 |  | h 9 g |  | 6.0 | 7.3 | 944 |  |
|  |  |  | h 8 e |  | 7.0 |  |  |  |
|  | 14.9 | I | h 8 g |  | 5.4 | 6.4 | 945 |  |
|  |  |  | h 7 e |  | 6.1 |  |  |  |
|  | 14.9 | II-III | h 8 g |  | 5.4 | 5.4 | 946 |  |
|  | 15.0 | DD | h 6 g |  | 4.0 | 4.0 | 953 | T very red |
|  | 15.0 | III | h 7 g ? |  | 4.5 | 4.5 | 959 | Fog |
|  | 15.2 | I | h 6 g g |  | 4.0 | 4.0 | 960 |  |
|  | 14.6 | I | h 6 g |  | 4.0 | " | 961 |  |
|  | 15.2 | I | h 6 g |  | 4.0 | " | 963 |  |
| Mar. $\quad 2$ | 15.5 | I | h 6 g |  | 4.0 | " | 968 |  |
| 5 | 14.8 | I | h 6 g |  | 4.0 | " | 971 |  |
| 25 | 14.4 | I | h 6 g |  | 4.0 | " | 09991 | $\left\{\begin{array}{l}\text { Near horizon } \\ \text { Not fainter }\end{array}\right.$ |

## METHOD BY STEPS:




1090
$\beta$ Persei (Algul) Serims V.
$(1900) 3^{\mathrm{h}} 1^{\mathrm{m}} 40^{\mathrm{s}}\left(+3^{\mathrm{s}} .89\right) ; \quad+40^{\circ} 34^{\prime} .2\left(+0^{\prime} .23\right)$
Period: $2^{d} 20^{\mathrm{h}} 48^{\mathrm{m}} \pm$; Variation: $2.3-3.5$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| $r$ Andr. | 11 | $+41^{\circ} 395$ | 0.0 | 2.3 HP. |
| $r$ Pers. | 19 | $52^{\circ} 654$ | 4.5 | 3.0 " |
| $\varepsilon$ | 32 | $39^{\circ} 895$ | 6.5 | 2.9 " |
| \% | 30 | $31^{\circ} 666$ | 7.0 | 2.9 " |
| i | 26 | $47^{\circ} 876$ | 9.3 | 3.1 " |
| $\rho$ | 20 | $38^{\circ} 630$ | $\left\{\begin{array}{l}11.6 \\ 12.1\end{array}\right.$ | 3.4-4.2 |
| $\kappa$ | 23 | $44^{\circ} 631$ | 12.6 | 4.1 HP . |
| , | 28 | $+42^{\circ} \mathrm{S} 15$ | 13.2 | 3.9 " |

## Notes:

As in the case of U Cephei, these observations were made mainly for practice between the regular observations of long period variables. For this reason the branches of the light curve observed are not symmetrical with regard to the minimum.

The comparison star $\rho$, which is irregularly variable, has two different steps assigned in the scale: 11.6 and 12.1, derived respectively from the observations on November 12 and April 12.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |

## METHOD BY STEPS:

| 87 Mar. | 22 | $14^{\mathrm{n}} 3^{\text {m }}$ |  | : $1 \beta 0$ | 7.3 |  | 7.3 | 10353 | Watch corr: $=-2^{\mathrm{m}}$ applied |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 1412 |  |  | 8.4 8.0 | 8.4 | 8.3 |  | Eph. Helioc. Min. $13^{\mathrm{a}} 27^{\mathrm{m}}$ |
|  |  | 1421 |  | $\varepsilon 1 \beta 2{ }^{\circ}$ | 7.4 | 7.4 | 7.3 |  |  |
|  |  |  |  | $\zeta 0$ \% | 7.0 |  |  |  |  |
|  |  | 1434 |  | ค18 | 5.5 |  | 5.9 |  |  |
|  |  |  |  | $\beta 1 \%$ | 6.0 |  |  |  |  |
|  |  | 1443 |  |  | 6.3 5.5 | 5.5 | 5.3 |  |  |
|  |  |  |  | ${ }_{\beta} 2{ }^{\text {\% }}$ | 5.0 |  |  |  |  |
|  |  | 1514 |  | \% A 4 ¢ $2 \varepsilon$ | 4.3 | 4.3 | - 4.2 |  | $\mathrm{A}=$ Andromedre |
|  |  | 161 |  | ¢ ${ }^{\text {¢ }}$ ¢ | 4.5 |  | 4.3 |  |  |
|  |  |  |  | P3: | 4.0 |  |  |  |  |
|  |  |  |  | F1\% $i 4 \%$ | 3.8 |  |  |  |  |
| 87 Nov. | 12 | 1426 | III | ¢ $2 \beta 2 \rho$ ? | 10.5 | 10.5 | 10.5 | 858 |  |
|  |  | 1439 |  | ${ }^{\circ} 1 \beta 2 \rho$ | 10.0 8 | 10.1 | 9.3 |  | Eph. Helioc. Min. $16^{\mathrm{b}} 17^{\mathrm{m}}$ |
|  |  | 1450 |  | ${ }^{\zeta} 1 \beta$ | 10.3 | 8.4 | 10.0 |  |  |
|  |  |  |  | ¢ $2 \beta 2 k$ | 9.8 | 9.8 |  | . |  |
|  |  | 154 |  | $\delta^{\circ} 1 \beta^{0} 0.5$ p | 10.7 | 10.8 | 10.0 |  |  |
|  |  |  |  | $\zeta 2 \beta 3 \kappa$ | 9.3 | 9.2 |  |  |  |
|  |  | $15 / 7$ |  |  | 10.7 9.3 | $\begin{array}{r} 10.8 \\ 9.2 \end{array}$ | 9.6 |  |  |
|  |  |  |  | $\varepsilon 1.5$ \% | 8.0 |  |  |  |  |
|  |  | 1530 |  |  | 11.5 |  | 10.1 |  |  |
|  |  |  |  | $\varepsilon 2.5 \beta$ | 9.0 |  |  |  |  |
|  |  |  |  | ¢ 2.5 B | 9.0 |  |  |  |  |
|  |  | 160 |  | ${ }_{\circ} 1 \beta 1 \rho$ | 10.5 | 10.5 | 9.7 |  |  |
|  |  |  |  | $\varepsilon 2 \beta$ | 8.5 |  |  |  | $\}^{\text {Hazy }}$ |
|  |  |  |  | $\varepsilon 1 \beta 2 \rho$ | 9.6 | 8.2 |  |  |  |
|  |  | 1630 |  |  | 9.0 |  | 7.9 |  |  |
|  |  |  |  | ¢ 0.5 \% | 7.0 |  |  |  |  |
| 88 Apr. | 12 |  | II-III |  |  |  |  |  |  |
|  |  | 1415 |  | $\varepsilon 0 \leftrightharpoons 1 \beta 3 \rho$ | 8.6 | 8.3 | 8.2 | 740 | Watch corr. $0^{\text {m }}$ |
|  |  | 1422 |  | ع $0<2 \beta 3 \nu$ | 9.2 | 9.3 | 9.5 |  |  |
|  |  |  |  | ¢1 1 ¢ 4 к | 9.5 | 10.0 |  |  |  |
|  |  |  |  |  |  |  |  |  |  |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| SS Apr. 12 | $14^{n} 41^{\text {m }}$ |  | $\varepsilon 3 \beta 3 \%$ | 9.9 | 9.8 | 10.3 | 10740 |  |
|  | 1450 |  | ع 3 \% 3 , | 9.9 | 9.8 | 10.3 |  |  |
|  | 157 |  | \% $2 \beta$ \% | 11.3 10.4 | 10.3 | 10.9 |  |  |
|  |  |  | ¢ 3 \% 1 \% | 11.7 | 11.1 |  |  |  |
|  | 1529 |  | i $3 \beta 3$, | 11.3 | 11.0 | 11.2 |  | Eph. Helioc. Min. $15^{\text {b }} 28^{\text {m }}$ |

s III.

Correction to :
21)

## "OBSERVATIONS OF VARIABLE STARS

made in the years 1884-1890, Washington 1901."
Page 22, Nov. 12,

$$
\begin{array}{cccccc}
\text { for } & 15^{\mathrm{h}} & 7^{\mathrm{m}} & \text { read } & 15^{\mathrm{h}} & 17^{\mathrm{m}} \\
\text { ، } & 15 & 14 & 6 & 15 & 44
\end{array}
$$

gn.

3
.5
9
3
5
1

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:




METHOD BY STEPS:


*) The original record has $\mathrm{R} 2 \mathrm{~b}, \mathrm{R} 3 \mathrm{f}$, contrary to the scale of steps.
(1900) $5^{\mathrm{h}} 9^{\mathrm{m}} 13^{\mathrm{s}}(+4.83) ;+53^{\circ} 28^{\prime} .4 \quad\left(+0^{\prime} .07\right)$

Period: 460.2 ; Variation: 7 ? $-12 \frac{1}{2}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps. | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| a | 1 | $+53^{\circ} 872$ | 0.0 | (6.5) BD . |
| g | 3 | 878 | 10.0 | 8.4 |
| c | 4 | 884 | 15.5 | 8.7 |
| b | 9 | 879 | 17.5 | 9.0 |
| d | 8 | $+53^{\circ} 880$ | 22.0 | 9.0 |
| e | 23 | ...... | 27.0 | 10.3 |
| f | 26 | $\cdots$ | 30.3 | 10.7 |

$\left.\begin{array}{l|c|c|c|c|c|c|c}\hline 1800+ & \text { Gr. M. T. } & \text { Sky } & \text { Comparisons } & \text { I } & \text { II } & \text { Mean } & 2400000\end{array}\right]$ Remarks

DECIMAL METHOD:



| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

METHOD BY STEPS:


Notes:
The observations of this star commenced soon after its discovery (1885) by Gore, and may for this reason be of value.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## METHOD BY STEPS :

| 86 Dec. | 7 | 15.5 |  | B 4 C 2 U 8 E | 3.5 | 4.9 | 4.2 | 10248 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 8 | 15.0 |  | B 3 C 5 U 8 E | 5.0 | 5.9 | 5.5 | 249 |
|  | 14 | 15.1 |  | B6C3U8E | 4.0 | 5.3 | 4.7 | 255 |
|  | 15 | 14.8 | I | B4C5U8E | 5.0 | 5.9 | 5.5 | 256 |
|  | 29 | 15.3 | I | B 2 C 3 U 8 E | 4.0 | 5.3 | 4.7 | 270 |
| 87 Jan. | 14 | 16.5 | III | B 5 U 4 k | 6.7 | 6.7 | 7.1 | 286 |
|  |  |  |  | B 5 U $7 \rho$ | 8.2 | 7.6 |  |  |
|  | 23 | 16.3 | I | B7 U3k | 8.2 | 8.6 | 8.8 | 295 |
|  |  |  |  | B $7 \mathrm{U} 7 \rho$ | 9.2 | 9.2 |  |  |
|  | 28 | 14.5 | II | B 8 U 4 k | 8.2 | 8.2 | 9.0 | 300 |
| Feb. | S | 15.8 |  | B $84 \mathrm{U} 7 \rho$ U 4 k | 9.7 | 9.8 | 8.8 |  |
|  |  |  |  | U9 9 | S.3 9.3 |  | 8.8 | 311 |
|  | 12 | 15.3 | I | U 4 k | 8.3 |  | 10.3 | 315 |
|  |  |  |  | U $6 \rho$ | 12.3 |  |  |  |
|  | 15 | 15.8 |  | k 1 U 5 ¢ | 13.3 | 13.3 | 13.3 | 318 |
|  | 16 | 14.7 | II | k1 U 5 P | 13.3 | 13.3 | 13.3 | 319 |
|  | 24 | 15.3 | I | k $2 \mathrm{U} 4 \rho$ | 14.3 | 14.3 | 14.3 | 327 |
|  | 27 | 16.1 | III | $\mathrm{k} 3 \mathrm{U} 4 \rho$ | 14.8 | 14.9 | 14.9 | 330 |
| Mar. | 13 | 15.8 | III | $\mathrm{k} 5 \mathrm{U} 2 \rho$ | 16.8 | 16.6 | 16.7 | 344 |
|  | 17 | 15.2 | I | k 4 U $2 \dot{\rho}$ | 16.3 | 16.3 | 16.3 | 348 |
|  | $\stackrel{27}{ }$ | 15.5 |  | \% 2 U 2 x | 20.3 | 20.3 | 20.3 | 358 |
|  | 28 | 16.2 | II | $p 2 \mathrm{U} 2 \mathrm{x}$ | 20.8 | 20.8 | 20.8 | 359 |
| Apr. | 18 | 15.1 | II | x 1 U 1 z ? | 24.6 | 24.6 | 24.6 | 380 |
|  | 20 | 14.8 | I | $x 2 \mathrm{U} 1 \mathrm{z}$ | 25.1 | 25.0 | 25.1 | 382 |
|  | 25 | 15.5 | I | x 2 U 0.5 z | 25.4 | 25.4 | 25.4 | 387 |
| 88 Jan. | 11 | 15.3 | I | E1 U 2 k | 10.3 | 10.2 | 11.0 | 648 |
|  |  |  |  | E1U5 ${ }^{\text {d }}$ | 11.8 | 11.6 |  |  |
|  | 17 | 15.2 | I D | E1 U 2 k | 10.3 | 10.2 | 11.0 | 654 |
| Feb. | 7 | 15.8 | I |  | 11.8 | 11.6 11.3 | 10.9 | 675 |
|  |  |  |  | B $7 \mathrm{U} 5 \rho$ | 10.2 | 10.7 | 10.9 | 670 |
|  | 14 | 15.8 | , | k 1 U 5 ¢ | 13.3 | 13.3 | 13.3 | 682 |
| Mar. | ${ }^{6}$ | 15.3 | I | k 3 U $2 \rho$ | 15.8 | 15.9 | 15.9 | 703 |
|  | 13 | 15.0 | I | k 3 U3 ${ }^{\text {a }}$ | 15.3 | 15.3 | 15.3 | 710 |
| Apr. | 3 11 | 15.5 | III | k5 U $0 \rho$ | 17.8 | 18.3 | 18.1 | 731 |
|  | 11 | 15.2 | I | $\rho 4 \mathrm{U} 2 \mathrm{x}$ | 21.8 | 21.6 | 21.7 | 739 |

## 2539 R Canis Minoris Series IV.

$(1900) 7^{\mathrm{h}} 3^{\mathrm{m}} 13^{\mathrm{s}}\left(+3 .{ }^{\mathrm{s}} 30\right) ;+10^{\circ} 10^{\prime} .9 \quad\left(-0^{\prime} .09\right)$ Period: $337^{\mathrm{d}} 7$; Variation; $7 \frac{1}{2}$. $-10^{\mathrm{m}}$.

Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | :--- |
|  |  |  |  |  |
| c |  | $+9^{\circ} 1539$ | 0.0 | 7.4 |
| a BD. |  | $10^{\circ} 1416$ | 6.7 | 8.4 |
| b |  | $10^{\circ} 1429$ | 11.0 | 8.3 |
| k |  | $9^{\circ} 1531$ | 15.0 | 8.9 |
| d |  | $10^{\circ} 1422$ | 18.7 | 9.0 |
| e |  | $10^{\circ} 1421$ | 24.7 | 9.5 |
| f |  | $+10^{\circ} 1418$ | 28.2 | 9.5 |
|  |  |  |  |  |

## Notes:

Two other comparison stars were used only once (1887, March 24), and are omitted in the scale and in the reductions. They are:

$$
\begin{aligned}
& \mathrm{g}=\mathrm{BD} .+10^{\circ} 1432, \stackrel{\mathrm{~m}}{9.0} \\
& \mathrm{~h}= \\
& \hline
\end{aligned}
$$

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | 2400000 | Remarks |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD :

|  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 83 Dec. | 21 | 15.6 | III | a 8 b a 7 b (Z) 7 | 10.1 9.7 | 9.9 | 09166 |
|  | 25 | 16.7 | III | c $7 \mathrm{~b}, \mathrm{R}>\mathrm{a}$ ! ! | 7.7 | 6.9 | 170 |
|  |  |  |  | c 5 b (Z) | 5.5 |  |  |
|  |  |  |  | a 2 b | 7.6 |  |  |
|  | 2 | 16.2 | I | $\mathrm{c}^{8} 8$ a | 5.4 | 6.1 | 173 |
| 84 Jan. | 3 | 15.4 | III | c 9 a | 6.0 | 6.4 | 179 |
|  |  |  |  | c 10 a (Z) | 6.7 |  |  |
|  | 5 | 15.3 | I D. | c 6 a | 4.0 | 3.4 | 181 |
|  |  |  |  | c 4 a (Z) | 2.7 |  |  |
|  | 15 | 14.7 | I | c 6 a | 4.0 | 4.0 | 191 |
|  | 17 | 14.2 | I |  | 4.0 2.7 | 2.4 | 193 |
|  |  |  | 1 | c 3 a (Z) | 2.0 | 2.4 |  |
|  | 19 | 15.5 | II | c 5 a | 3.4 | 3.7 | 195 |
|  |  |  |  | c 6 a (Z) | 4.0 |  |  |
|  | 20 | 14.3 | II | c 4 a | 2.7 | 2.7 | 196 |
|  |  |  |  | c 4 a (Z) | 2.7 |  |  |
|  | 21 | 15.1 | II | c 3 a | 2.0 | 1.7 | 197 |
|  |  |  |  | c 2 a (Z) | 1.3 |  |  |
|  | 23 | 14.3 | I | c 3 a | 2.0 | 1.7 | 199 |
|  |  |  |  | c 2 a (Z) | 1.3 |  |  |
|  | 24 | 15.1 | II | c 3 a | 2.0 | 2.4 | 200 |
|  |  |  |  | c 4 a (Z) | 2.7 |  | 206 |
|  | 30 | 14.5 | II | c 3 a ${ }_{\text {c }}$ (Z) | 2.7 2.0 | 2.4 | 206 |
|  | 31 | 15.3 | III | c 3 a | 2.0 | 2.4 | 207 |
|  |  |  |  | c 4 a (Z) | 2.7 |  |  |
| Feb. | 14 | 15.1 | III | c 3 a | 2.0 | 2.4 | 221 |
|  |  |  |  | c 4 a (Z) | 2.7 |  |  |
|  | 19 | 15.4 | I | c 3 a ${ }_{\text {c }}$ a $(\mathrm{Z})$ | 2.0 1.3 | 1.7 | 226 |
| Mar. | 23 | 15.1 | III | c 7 a | 4.5 | 4.5 | 259 |
|  |  |  |  | c 7 a (Z) | 4.5 |  |  |
| Apr. | 2 | 15.8 | I D | c 8 a | 5.4 | 4.7 | 269 |
|  |  |  |  | c 6 a (Z) | 4.0 |  |  |
|  | 20 | 14.6 | III |  | 11.8 | 13.0 | 287 |
| May | 13 | 15.3 | II | b 4 d $(\mathrm{Z})$ R | 14.1 | > 19 | 310 |
| Nov. | 18 | 16.1 | I | a 8 b | 10.1 | 11.4 | 499 |
|  |  |  |  | a 5 d | 12.7 |  |  |
| Dec. | 22 | 15.5 | II | c 6 a | 4.9 | 3.7 | 533 |
|  |  |  |  | c 3 b | 3.3 |  |  |



METHOD BY STEPS:


*) See notes above.


Period: $305^{\mathrm{A}} .0$; Variation: $9 \frac{1}{2}^{\mathrm{N}}-<14^{\mathrm{n}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
|  |  |  |  |  |
| b | 6 | $+19^{\circ} 2045$ | 0 |  |
| c | 12 | 2046 | 9 | 8.9 |
| d | 20 | 2048 | 15 | 9.3 |
|  |  |  |  | 9.9 |

Notes:
This variable proved to be too faint for the 3 -inch glass. When it was marked "barely visible," it must has been between the 10th and 11th magnitude. The few observations may serve to confirm those made elsewher

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |



$$
\begin{aligned}
& 3109 \\
& \text { S Cancri }
\end{aligned} \text { Series IV. }
$$

Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  | $+19^{\circ} 2097$ | 0.0 |  |
| b |  | 2094 | 8.0 | 8.5 " |
| d |  | 2101 | 9.6 | 8.7 " |
| c |  | 2088 | 12.9 | 9.0 " |
| e |  | 2089 | 21.8 | 9.4 " |
| k |  | 2086 | 25.7 | 9.5 " |
| g | , | $+19^{\circ} 2085$ | 29.6 | 9.5 " |

## Notes:

The observations of this Algol-Star were made partly for the sake of practice, partly with the view of studying its general light curve. While pursuing them a secondary minimum was suspected.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:





METHOD BY STEPS :

| 87 Jan. 23 | $\begin{aligned} & 1426 \\ & 1531 \\ & 1614 \end{aligned}$ | I I II |  | $\begin{aligned} & 24.3 \\ & (22.6) \\ & 21.3 \\ & 20.8 \\ & 25.6 \\ & 21.3 \\ & 19.8 \\ & 22.6 \end{aligned}$ | 24.1 <br> 21.9 <br> 21.4 | 23.8 <br> 22.4 <br> 21.3 | 295 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

3477 R Leonis Minoris Series III.
(1900) $9^{\mathrm{h}} 39^{\mathrm{m}} 35^{\mathrm{s}}(+3.61) ;+34^{\circ} 58^{\prime} .3 \quad\left(-0^{\prime} .27\right)$

Period: $370^{\circ} 5 \pm ; \quad$ Variation: $7^{\mathrm{x}}-13^{\mathrm{m}}$
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| B | 1 | $+35^{\circ} 2042$ | 0.0 | (6.5) BD. |
| C | 2 | +34 ${ }^{\circ} 2035$ | 3.8 | (6.3) " |
| E | 3 | +34 ${ }^{\circ} 2022$ | 5.5 | 7.6 |
| H | 4 | + $35^{\circ} 2046$ | 9.5 | 7.9 |

Notes:
This variable was observed for a short while at the request of Mr. H. M. Parkhurst. The comparison stars and their designations are identical with those published by him in the Annals of H. C. O., vol. XXIX, page 150

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |



3825 . R Ursæ Maioris Series III.
(1900) $10^{\mathrm{h}} 37^{\mathrm{m}} 34^{\mathrm{s}}\left(+4 .{ }^{\mathrm{s}} 32\right) ; \quad+69^{\circ} 18^{\prime} .0 \quad\left(-0^{\prime} .31\right)$

Period: $302^{\mathrm{d}} 1 \pm ; \quad$ Variation: $7^{\mathrm{M}}-13^{\mathrm{m}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| a | 1 | $+69^{\circ} 586$ | 0.0 | (4.7) BD. |
| f | 2 | 583 | 2.1 | (5.5) " |
| b | 5 | 584 | 10.2 | 8.5 |
| c | 7 | $+69^{\circ} 585$ | 11.9 | 9.1 |
| d | 12 | . | 14.0 | 9.9 |
| e | 14 | ….. | 16.0 | 10.6 |

## Notes:

As the observations of this star were discontinued before Argelander's method by steps was employed they furnish ro scale for the comparison stars. An artificial scale was constructed from the data of the ASV The scale of the Atlas does not comprise the stars $a$ and $f$, but can be extended by extrapolation by means of th formula which connects steps and magnitude. This gives the steps - 88 and -67 for $a$ and $f$ respectively. By adding +88 to all the numbers of the scale and finally dividing by 10 the above scale was found.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:




5157 S Bootis Series III.
(1900) $14^{\mathrm{n}} 19^{\mathrm{m}} 32^{\mathrm{s}}\left(+2^{\mathrm{s}} .01\right) ;+54^{\circ} 15^{\prime} .9$ ( $-0^{\prime} .27$ )

Period: $268^{\mathrm{d}} .2+; \quad$ Variation: $8^{\mathrm{M}}-13^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| ---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
|  |  |  |  |  |
| k | 2 | $+54^{\circ} 1668$ | 0.0 | 8.1 |
| p | 3 | $+53^{\circ} 1714$ | 2.6 | 8.2 |
| g | 4 | $+54^{\circ} 1679$ | 5.5 | 8.5 |
| r | 7 | 1677 | 12.0 | 9.3 |
| e | 8 | 1676 | 15.3 | 9.5 |
| c | 9 | 1663 | 16.5 | 9.5 |
| a | 14 | 1674 | 21.3 | 9.9 |
| d | 12 | 1672 | 22.8 | 9.8 |
| m | 15 | $+54^{\circ} 1670$ | 24.3 | 10.1 |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD :



METHOD BY STEPS:


5484 U Coronæ $\quad$ Series IV.

Period: $3^{\mathrm{d}} 10^{\mathrm{h}} 51^{\mathrm{m}} 12 . .^{\text {s. }} 4-$; Variation: $7^{\mathrm{m}} 5-8 .{ }^{\mathrm{m}} .9$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | :---: | :---: | :---: |
|  |  |  |  |  |
| a | III 2 | $+31^{\circ} 2724$ | 0.0 | 7.6 |
| h | - | $+32^{\circ} 2575$ | $(8.2)$ | $(8.1) \mathrm{BD}$. |
| g | 4 | $+32^{\circ} 2578$ | 8.4 | 8.1 |
| b | 5 | $+32^{\circ} 2577$ | $(11.0)$ | 8.4 |
| i | - | $+32^{\circ} 2573$ | 11.2 | $(8.9) \mathrm{BD}$. |
| c | 6 | $+32^{\circ} 2572$ | 16.2 | 8.6 |
| d | 10 | $\ldots$ | $(22.9)$ | 9.8 |

## Notes:

This variable is on the Chart for S Coronæ ( 5504 , Series III), and the numbers in the 2 d and 5 th column refer to the Catalogue of the same variable. The steps in parentheses could not be determined with accuracy That of the comparison star $d$, which occurs only once, is taken from the scale for S Coronæ.

The observations, like all those of the Algol type variables, are not systematically arranged for the determina tion of the minima.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I. | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:

| 84 June July | 28 12 | $16^{\text {b }} 0^{\text {m }}$ 1510 | II II | a 8 c a 6 c a a c c | 13.0 9.7 16.2 | 11.4 14.6 | 09356 370 | ${ }^{\frac{1}{2}}{ }^{\text {a }}$ before Eph. Min. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| July |  |  |  | a $8 \mathrm{c}(\mathrm{Z})$ | 13.0 |  | 370 |  |
|  | 18 | 1532 | III | a 4 c | 6.5 | 7.3 | 376 |  |
|  | 20 | 1532 |  | a 5 c (Z) | 8.1 |  |  |  |
|  | 20 | 1032 | II | a 4 c a $4 \mathrm{c}(\mathrm{Z})$ | 6.5 6.5 | 6.5 | 378 |  |
| Aug. | 21 | 1520 | I | a 5 c | 8.1 | 7.3 | 410 |  |
| Sept. | 9 | 1443 | I | a $4 \mathrm{c}(\mathrm{Z})$ | 6.5 | $5.4$ | 429 | $1^{\text {b }}$ before Eph. Min. |
|  |  |  |  | a 3 c | 4.9 |  |  |  |
| Oct. | 19 | 1515 | II | a $\begin{aligned} & \text { c } \\ & \text { d }\end{aligned}$ | 5.9 16.9 | 16.9 | 439 |  |
|  | 9 | 1420 | II | a 2 c a 7 g | 3.2 5.9 | 4.6 | 459 |  |
| 85Apr.May | 10 | 1613 | II | a 4 c | 6.5 | 5.8 | 642 |  |
|  | 3 | 1647 | III | a 6 g | 5.0 |  |  |  |
|  |  | 164 | III | a 4 c | 6.5 6.6 | 6.6 | 665 |  |
|  | 12 | 1617 | I | a 4 c g 1 b | 6.5 | 7.6 | 674 |  |
| June | 4 | 1647 | I | a 4 ca 7 g | 8.7 6.5 |  | 697 |  |
|  |  | 1552 |  |  | 5.9 | 6.2 |  |  |
|  | 16 |  | I | a 5 c | 8.1 | 8.4 | 709 |  |
| July |  | 1620 |  | g 1 b | 8.7 |  |  |  |
|  | 1. |  | I | ${ }_{\text {a }} 4 \mathrm{c}$ | 6.5 | 7.9 | 724 |  |
|  |  |  |  | g 1 c | 9.2 |  |  |  |
|  | 13 | 1615 | I | a 6 c | 9.7 | 9.7 | 736 |  |
| Aug. | 30 | 1542 | I |  | 15.7 | 15.7 | 760 | $2^{\text {n }}$ before Eplı. Min. |
|  |  | 1528 | I | b 9 c a 6 b | 6.6 | 15.7 5.8 | 784 |  |
| Oct. |  |  |  | a 3 c | 4.9 |  |  |  |
|  | 5 | 1442 | I | a 4 c | 6.5 | 7.3 | 820 |  |
|  |  |  |  | a 7 b | 7.7 |  |  |  |
|  |  |  |  | a 9 g | 7.6 |  |  |  |
| 86 Mar. | 9 | 1542 | I | a 4 ca 5 b | 6.5 | 6.0 | 971 | Near hor. |
|  | 22 | 1432 |  |  | 5.5 | 12.3 | 988 |  |
|  |  |  | I | a 8 c | 13.0 11.5 |  |  | Eph. Hel. Min. $14^{\text {b }} 27^{\text {a }}$ |
|  | " | 1439 | " | a 7 c | 11.3 | 11.3 | " |  |
|  | " | 1445 | " | a 7 c | 11.3 | 11.4 | " |  |
|  |  | 1452 |  | b 1 c | 11.5 |  |  |  |
|  | " |  | " | a 7 c | 11.3 | 11.1 | " |  |
|  |  |  |  | g 10 b | 11.0 |  |  |  |
|  |  |  |  | g 4 c | 11.5 |  |  |  |
|  |  |  |  | g. 9 b | 10.7 |  |  |  |


| 1800 |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 86 Mar. | 22 | $15^{\mathrm{n}} \quad 0^{\text {m }}$ | I | a 7 c |  | 11.3 | 11.1 | 09988 |  |
|  |  |  |  | g 8 b |  | 10.5 |  |  |  |
|  |  | 159 | " | g 46 a 6 c |  | 11.5 9.7 | 10.1 | " |  |
|  |  |  |  | g 6 b |  | 10.0 |  |  |  |
|  |  |  |  | g 3 c |  | 10.7 |  |  |  |
|  |  | 1520 | " | a 6 c |  | 9.7 | 9.9 | " |  |
|  |  |  |  | g 6 b |  | 10.0 |  |  |  |
| Apr. | ${ }_{5}^{1}$ | 1544 1450 | $\stackrel{\text { I }}{\text { II }}$ | a 7 c a 4 c |  | 11.3 6.5 | 11.3 6.6 | 09998 10902 |  |
|  |  |  |  | a 8 g |  | 6.7 |  |  |  |
|  | 22 | 1530 | III | a 6.5 c |  | 10.5 | 10.6 | 019 | $\frac{1}{2}{ }^{\text {n }}$ before Eph. Min. |
|  |  |  |  | h 3 c |  | 10.6 |  |  |  |
| May | 6 | 1453 | III | a 6 c |  | 9.7 | 8.7 | 033 |  |
|  |  |  |  | a 9 g |  | 7.6 |  |  |  |
|  |  | 1534 | " | a 8 g |  | 6.7 | 8.2 | " |  |
| June | 30 | 1537 | II-III | a 6 c a 8 c |  | 9.7 13.0 | 12.6 | 088 |  |
|  |  |  |  | i 2 c |  | 12.2 |  |  |  |

## METHOD BY STEPS:

| 87 May | 24 | 1542 |  | j 1 U | 12.2 |  | 12.2 | 416 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 1552 |  | i 1 U 3 c | 12.7 | 12.4 | 12.6 | " |  |
|  |  | 1615 |  | i 1 U 3 c | 12.7 | 12.4 | 12.6 | " | Eph. Hel. Min. $16^{\text {b }} 21^{\text {m }}$ |
|  |  | 1629 |  | i 2 U 2 c | 13.7 | 13.7 | 13.7 | " |  |
|  |  | 179 |  | i1 U 3 c | 12.7 | 12.4 | 12.6 | " |  |
| July | 9 | 150 | II | U 3 i | 8.2 |  | 8.2 | 462 |  |
|  | 15 | 1512 | I | U3i | 8.2 |  | 8.2 | 468 |  |
| Sept. | 15 | 1444 | I | i 1 U 3 c | 12.7 | 12.4 | 12.6 | 530 | Eph. Hel. Min. $14^{\text {b }} 32^{\text {m }}$ |
|  |  | 1457 | " | i 2 U 3 c | 13.2 | 13.2 | 13.2 | " |  |
|  |  | 154 | " | g 2 U 0 i | 10.8 |  | 10.8 | " |  |
|  |  | 1513 | " | g 2 U 1 i | 10.3 | 10.3 | 10.3 | " |  |
|  |  | 1524 | " | g 2 U 1 i | 10.3 | 10.3 | 10.3 | " |  |
|  |  | 1536 | " | g 1 U 2 i | 9.3 | 9.3 | 9.3 | " |  |
| 88 Sept. | 8 | 156 | I | i 2 U 3 c | 13.2 | 13.2 | 13.2 | 889 |  |
| - |  | 1514 | " | i 3 U 3 c | 13.7 | 13.7 | 13.7 | " |  |
|  |  | 1525 | " | i 3 U 3 c | 13.7 | 13.7 | 13.7 | " | Eph. Hel. Min. $15^{\mathrm{h}} 21^{\mathrm{m}}$ |
|  |  | 1537 | " | i 2 U 3 c | 13.2 | 13.2 | 13.2 | " |  |
|  |  | 1544 | " | i 2 U 4 c | 12.7 | 12.9 | 12.8 | " |  |
|  |  | 1556 | " | i 1 U 5 c | 11.7 | 12.0 | 11.9 | " |  |
|  |  | 160 | " | i 0 U | 11.2 |  | 11.2 | " |  |

5501 S Serpentis Series II.
(1900) $15^{\mathrm{h}} 16^{\mathrm{m}} 59^{\mathrm{s}}\left(+2^{\mathrm{s}} .81\right) ; \quad+14^{\circ} 40^{\prime} .4 \quad\left(-0^{\prime} .22\right)$

Period: $365^{\mathrm{d}} 4$; Variation: $8^{\mathrm{m}}-12 \frac{1}{2}^{\mathrm{m}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
| b | 1 | $+15^{\circ} 2845$ | 0.0 | 8.0 |
| a | 2 | $14^{\circ} 2866$ | 8.5 | 8.3 |
| c | 3 | $14^{\circ} 2868$ | 11.2 | 8.6 |
| e | 4 | $15^{\circ} 2846$ | 15.0 | 9.0 |
| d | 5 | $14^{\circ} 2862$ | 19.0 | 9.3 |
| f | 6 | $+15^{\circ} 2848$ | 20.4 | 9.5 |
| g | 8 | $\ldots \ldots \ldots$ | 24.0 | 9.9 |

- Notes:

The comparison star $f$, which was not used with the method by steps, was inserted in the above scale by means of observations made later for the Atlas.


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | .II | Mean | $2400000+$ | Remarks |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

METHOD BY STEPS:

| 87 Mar. | 24 | 16.5 | I | a 2 S 1 c | 10.4 | 10.3 | 10.4 | 10355 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Apr. | 29 | 15.7 | III | c 2 S 2 e | 13.1 | 13.1 | 13.1 | 391 |  |
| May | 15 | 15.8 | I | e 1 S 1 d | 17.0 | 17.0 | 17.0 | 407 |  |
|  | 20 | 15.5 | III | e 2 Sod | 18.0 |  | 18.0 | 412 |  |
| June | 14 | 15.9 | I | d 2 S 2 g | 21.5 | 21.5 | 21.5 | 437 |  |
|  | 17 | 16.6 | I | d 4 S 2 g | 22.5 | 22.3 | 22.4 | 440 |  |
|  | 23 | 15.5 | I | $\mathrm{S}=\mathrm{g}!$ ! | 24.0 |  | 24.0 | 446 |  |
| July | 9 | 15.5 | II | g 2 S | 26.0 |  | 26.0 | 462 |  |
| 88 Apr. | 5 | 15.6 | III | b 3 S 4 a | 3.8 | 3.6 | 3.7 | 733 |  |
|  | 11 | 15.4 | I | b 4 S 6 a | 3.3 | 3.4 | 3.4 | 740 |  |
| May | 28 | 15.5 | I | a 5 S 5 d | 13.8 | 13.8 | 14.8 | 786 |  |
| June | 2 | 15.5 | I | e 2 S 5 d a 4 S 4 d | 15.5 13.8 | 16.1 13.8 | 14.7 | 791 | c 2 d |
|  |  |  |  | e 1S4d | 15.5 | 15.8 |  |  |  |
|  | 29 | 15.2 | I | d 2 S 3 g | 21.0 | 21.0 | 21.0 | 818 |  |
| July | 11 | 16.1 | $\stackrel{\text { I }}{\text { I }}$ | d 5S 3 g | 22.5 | 22.1 | 22.3 | 830 |  |
|  | 29 | 16.3 | III | g 2 S | 26.0 |  | 26.0 | 848 | Seeing poor. |

*) The journal has a 5 S 3 g ; but it is evident from the scale of the comparison stars, that $a$ would never $b$ combined with $g$ as long as $d$ was available.
5504 S Coronæ Serirs III.
(1900) $15^{\mathrm{h}} 17^{\mathrm{m}} 19^{\mathrm{s}} \quad\left(+2^{\mathrm{s}} .45\right) ; \quad+31^{\circ} 43^{\prime} .6 \quad\left(-0^{\prime} .22\right)$

Period: 360.8; Variation: $7^{\mathrm{x}}-12^{\mathrm{x}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| f | - | $+33^{\circ} 2574$ | 0.0 | (6.8) BD. |
| a | 2 | $31^{\circ} 2724$ | 2.0 | 7.6 |
| g | 4 | $32^{\circ} 2578$ | 7.1 | 8.1 |
| $h_{1}$ | - | $32^{2} 2575$ | 9.0 | (8.1) BD. |
| b | 5 | $32^{\circ} 2577$ | 11.1 | 8.4 |
| c | 6 | $+32^{\circ} 2572$ | 16.1 | 8.6 |
| d | 10 | - ........ | 22.9 | 9.8 |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD :




METHOD BY STEPS:

| 87 Mar. | 17 | 15.5 | I | a 2 S 2 g a 2 S 5 b | 4.6 5.1 | 4.6 4.6 | 1.7 | 348 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 24 | 16.1 | I | a 2 S 3 g | 4.1 | 4.0 | 4.1 | 355 |  |
| Apr. | 20 | 15.0 | I | S 0 a | 2.0 |  | 2.0 | 382 |  |
|  | 25 | 16.4 | I | a 1 S | 3.0 |  | 3.0 | 387 |  |
| May | 15 | 16.0 | I | a 2 S 3 g ! | 4.1 | 4.0 | 4.1 | 407 |  |
|  | 20 | 15.7 | III | a 4 S 2 g ! | 5.6 | 5.4 | 5.5 | 412 |  |
|  | 24 | 15.8 |  | a 3 S 2 g | 5.1 | 5.1 | 5.1 | 416 |  |
| June | 14 | 16.0 | I | g 1 S 2 b | 8.6 | 8.4 | - 8.5 | 437 |  |
|  | 17 | 16.7 | I | g 1 S 2 b | 8.6 | 8.4 | 8.5 | 440 |  |
|  | 23 | 15.7 | I | g 3 S 2 b | 9.6 | 9.5 | 9.6 | 446 |  |
| July | 9 | $\cdot 15.7$ | II | b 2 S 5 d | 15.5 | 14.5 | 15.0 | 462 |  |
|  | 15 | 15.2 | I | b 4S 3d | 17.5 | 17.8 | 17.5 | 468 |  |
|  | 22 | 15.5 | I | b ${ }_{\text {c }}$ d S 4 d | 17.0 | 17.0 | 16.7 | 475 |  |
|  |  |  |  | c 0 S | 16.1 |  |  |  |  |
| Aug. | 6 | 15.0 | II | c 3 S 3 d | 19.5 | 19.5 | 19.5 | 490 |  |
| Sept. | 6 | 15.3 | II | d 2 S | 24.9 |  | 24.9 | 521 |  |
|  | 13 | 14.7 | I | S 1 d ! | 21.9 |  | 21.9 | 528 |  |
|  | 15 | 15.6 | I | d 1 S | 23.9 |  | 23.9 | 530 |  |
| 88 Apr. | 5 | 15.5 | II-III | S 0 a | 2.0 |  | 2.0 | 733 |  |
| - | 11 | 15.3 | I | f 0 S 2 a | 0.0 |  | 0.0 | 739 |  |
| May | 28 | 15.6 | I | a 3 S 3 g | 4.6 | 4.6 | 4.6 | 786 |  |
| June | 2 | 15.6 | I | a 3 S 4 b | 6.1 | 5.9 | 6.0 | 791 |  |
|  | 29 | 15.4 | I | g 1 S 2 h | 7.6 | 7.7 | 7.5 | 818 |  |
| July | 11 | 16.2 | I | g 2 S 3 b | 8.6 | 8.7 | 8.8 | 830 |  |
|  |  |  |  | S 0 h | 9.0 |  |  |  |  |
|  | 29 | 16.4 | I | b 4S 2 c | 14.6 | 14.4 | 14.5 | 848 | Seeing poor. |
| Aug. | 7 | 15.3 | I | b 6 S 6 d | 17.0 | 17.0 | 17.3 | 857 |  |
|  |  |  |  | c 2 S 6 d | 17.5 | 17.8 |  |  |  |
|  | 12 | 15.4 | I | c 3 S 4 d | 19.0 | 19.0 | 19.0 | 862 |  |
|  | 24 | 15.2 | I | c 5 S 3 d | 20.5 | 20.4 | 20.5 | 874 |  |
|  | 29 | 15.3 | I | c 4 S 3 d | 20.0 | 20.0 | 20.0 | 879 |  |
| Sept. | 6 | 14.9 | I | c 7 S 3 d | 21.5 | 20.9 | 21.2 | 887 |  |


| 5770 | R Herculis | Series II. |
| :---: | :---: | :---: |
| $(1900) 16^{\mathrm{h}} 1^{\mathrm{m}} 44^{\mathrm{s}}\left(+2^{\mathrm{s}} .68\right) ;+18^{\circ} 38^{\prime} .4$ | $\left(-0^{\prime} .16\right)$ |  | Period: $317 .{ }^{\mathrm{A}} 7 \pm$; Variation : $8 \frac{1^{\mathrm{N}}}{}-<13^{\mathrm{M}}$.

Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
|  |  |  |  |  |
| f | 3 | $+18^{\circ} 3113$ | 0.0 | 8.6 |
| e | 4 | 3114 | 4.0 | 8.7 |
| c | 5 | 3120 | 6.0 | 8.9 |
| d | 7 | 3121 | 8.7 | 9.1 |
| a | 9 | 3119 | 10.9 | 9.5 |
| g | 10 | 3111 | 11.5 | 9.6 |
| k | 12 | 3115 | 14.5 | 9.8 |
| b | 18 | $+18^{\circ} 3116$ | 17.5 | 10.4 |

Notes:
Comparison star $a$, which was used only once, has been inserted from later observations with the 12 -in refractor at Georgetown.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:



| Gr. M. T. | Sky | Comparisons | I | II | Mean | 2400000+ | Remarks |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 86 July | 4 | 16.3 | I | d 6 g <br> c 8 g |  | 10.5 <br> 10.5 | 10.5 | 10092 |


| 87 Apr. | 20 | 15.3 | I | c 2 R 1 d | 7.9 | 7.8 | 7.9 | 382 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 25 | 16.6 | I | c 0 R 2 d | 6.4 |  | 6.4 | 387 |
| May | 18 | 16.2 | I | d 2 R 1 g | 10.6 | 10.6 | 10.6 | 410 |
|  | 24 | 16.2 |  | g $1 \mathrm{R}_{\mathrm{R}} 2 \mathrm{k}$ | 12.5 | 12.5 | 12.2 | 416 |
| June | 15 | 15.6 | I | b 1 R | 18.5 |  | 18.5 | 438 |
|  |  |  |  | k 4 R | 18.5 |  |  |  |
|  | 19 |  | I | b 2 R | 19.5 |  | 19.5 | 442 |
| 88 Apr. | 5 | 16.0 | II-III | f 4 R 2 c R 3 d | 4.0 5.7 | 4.0 | 4.4 | 733 |
|  |  |  |  | R 0 e | 4.0 |  |  |  |
|  | 11 | 15.6 | I | c 1 R 2 d | 6.9 | 6.8 | 7.2 | 739 |
|  |  |  |  | e 4 R R invis. | 8.0 |  |  |  |
| June | 30 | 16.3 | II | "" " |  |  | > ${ }^{\text {c }}$ | 819 |

5950
W Herculis Serifs III.
(1900) $16^{\mathrm{h}} 31^{\mathrm{m}} 41^{\mathrm{s}}\left(+2^{\mathrm{s}} .13\right) ;+37^{\circ} 32^{\prime} .4 \quad\left(-0^{\prime} .13\right)$

Period: 280 $.0 \pm$; Variation: $8^{\mathrm{n}}-<13^{\mathrm{M}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| g | 2 | $+37^{\circ} 2774$ | 0.0 | 8.2 |
| a | 4 | $37^{\circ} 2772$ | 2.5 | 8.4 |
| f | 5 | $38^{\circ} 2801$ | 5.0 | 8.4 |
| c | 9 | $37^{\circ} 2775$ | 7.5 | 9.0 |
| b | 14 | $+37^{\circ} 2773$ | 8.5 | 9.5 |
| e | 18 | -....... | 12.0 | 10.0 |
| d | 19 | ........ | 16.5 | 10.3 |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |

## DECIMAL METHOD :



## METHOD BY STEPS:



| $1800+$ |  |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Oct. | 11 18 | 15.8 14.5 | $\stackrel{\mathrm{I}}{\mathrm{II}}$ | $\begin{aligned} & \mathrm{e} 2 \mathrm{~W} \quad 2 \mathrm{~d} \\ & \text { e } 1 \mathrm{~W} 4 \mathrm{~d} \\ & \mathrm{~b} 3 \mathrm{~W} \end{aligned}$ | $\begin{aligned} & 14.3 \\ & 12.8 \\ & 11.5 \end{aligned}$ | $\begin{aligned} & 14.3 \\ & 12.9 \end{aligned}$ | $\begin{aligned} & 14.3 \\ & 12.4 \end{aligned}$ | $10 \quad 556$ |  |
| 88 | Apr. <br> May <br> June | $\begin{array}{r} 5 \\ 28 \\ 30 \end{array}$ | $\begin{aligned} & 16.2 \\ & 15.9 \\ & 16.4 \end{aligned}$ | $\begin{aligned} & \text { III } \\ & \text { I } \\ & \text { II } \end{aligned}$ | W invis. |  |  | $>17$ $"$ " | $\begin{aligned} & 733 \\ & 786 \\ & 819 \end{aligned}$ |  |

6044 S Herculis Series II.
$(1900) 16^{\mathrm{h}} 47^{\mathrm{m}} 21^{\mathrm{s}} \quad\left(+2^{\mathrm{s}} .73\right) ;+15^{\circ} 6^{\prime} .6 \quad\left(-0^{\prime} .10\right)$
Period: $308^{\mathrm{d}} .1$; large irregularities; Variation: $7^{\mathrm{M}}-12^{\mathrm{M}}$
Comparison Stars:

| Obs. | ASV. | BD . | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| a | 1 | $+15^{\circ} 3066$ | 0.0 | (6.1) BD. |
| c | 5 | 3070 | 7.5 | 8.6 |
| b | 6 | 3060 | 10.0 | 8.8 |
| d | 13 | $+15^{\circ} 3062$ | 15.8 | 9.8 |


| $1800+$ | Gr. M.T. Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:




METHOD BY STEPS:

| 87 | Apr. | 20 | 15.7 | I | b 3 S 2 d | 13.4 | 13.5 | 13.5 | 382 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 29 | 16.0 | III | d 1 S | 16.8 |  | 16.8 | 391 |
|  | May | 18 | 16.5 | I | d 2 S | 17.8 |  | 17.8 | 410 |
|  | Sept. | 6 | 15.5 | II | S invis. |  |  | > 16 | 521 |
| 88 | Apr. | 5 | 16.4 | III | S invis. |  |  | $>16$ |  |
|  | May | 28 | 15.9 | I | " |  |  | " | 786 |
|  | June | 30 | 16.5 | II | " |  |  | . | 819 |

T Herculis Series III
(1900) $18^{\mathrm{h}} 5^{\mathrm{m}} 19^{\mathrm{s}}\left(+2^{\mathrm{n}} .27\right) ;+31^{\circ} 0^{\prime} .2 \quad\left(+0^{\prime} .01\right)$

Period: $164^{a} .85 \pm ;$ Variation: $8^{\mathrm{N}}-11 \frac{1}{2}^{\mathrm{m}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | :---: | :---: | :---: |
|  |  |  |  |  |
| g | 3 | $+30^{\circ} 3138$ | 0.0 |  |
| e | 4 | $+30^{\circ} 3133$ | 4.0 | 7.5 |
| a | 6 | $+30^{\circ} 3142$ | 8.0 | 7.9 |
| f | 14 | $+30^{\circ} 3139$ | 16.0 | 8.1 |
| b | 22 | $+30^{\circ} 3136$ | 19.7 | 9.1 |
| d | 29 | $+31^{\circ} 3185$ | 23.8 | 9.6 |
| c | 31 | $+30^{\circ} 3135$ | 27.8 | 10.1 |
|  |  |  | 10.2 |  |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD :



| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{array}{rr} 86 & \text { June } \\ \text { July } & 23 \\ 1 \end{array}$ | 17.3 15.5 15.7 | $\begin{aligned} & \mathrm{I} \\ & \mathrm{II} \\ & \text { III } \end{aligned}$ | $\mathrm{T}<\mathrm{d}$ <br> T barely vis. T invis. |  |  | $>28$ <br> $"$ <br>  | $\begin{array}{r} 10060 \\ 081 \\ \quad 089 \end{array}$ |  |

METHOD BY STEPS:


Comparison Stars:

| Obs. | ASV. | BD . | Steps. | Magn |
| :---: | :---: | :---: | :---: | :---: |
| d | 2 | $+49^{\circ} 3059$ | 0.0 | (7.0) |
| f | 5 | 3051 | 9.3 | 8.6 |
| c | 6 | 3073 | 11.3 | 8.8 |
| a | 11 | 3072 | 12.8 | 9.1 |
| b | 14 | 3065 | 15.0 | 9.3 |
| e | 31 | $+49^{\circ} 3068$ | 18.0 | 10.2 |



DECIMAL METHOD :

| $\begin{aligned} & 83 \text { July } \\ & \text { Aug. } \end{aligned}$ | 31 | 16.4 | I | R barely vis. |  | $>18$ | 09023 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 2 | 16.7 | I | " " |  | > | 025 |
|  | 4 | 16.4 | II | " " |  | " | 027 |
|  | 22 | 15.1 | I | " |  | " | 045 |
| Sept. | 2. | 15.3 | I | " " |  | " | 056 |
| Oct. | 26 | 15.8 | 1 | R invis. |  | " | 080 |
|  | 30 | 14.6 | I | R barely vis. |  | " | 084 |
|  | 6 | 15.8 | I |  |  | " | 090 |
|  | 21 | 15.5 | ? | a 6 b | 14.1 | 14.1 | 105 |
|  | 29 | 15.0 | I | c 8 a | 12.5 | 12.5 | 113 |
| Nov. | 1 | 15.7 | I | d 6 c | 6.8 | 6.8 | 116 |
|  | 3 | 14.6 | I | d 4.5 c | 5.1 | 5.1 | 118 |
|  | 6 | 14.8 | D | d 4 c | 4.5 | 4.5 | 121 |
|  | 7 | 14.4 | 2 | d 3 c | 3.4 | 3.4 | 122 |
|  | 18 | 14.4 | II | d 2.5 c | 2.8 | 2.8 | 133 |
|  | 26 | 15.0 | I | d 2 c | 2.3 | 2.3 | 141 |
|  | 27 | 14.9 | III | d 1 c | 1.1 | 2.0 | 142 |
|  |  |  |  | d 2.5 c (Z) | 2.8 |  |  |
|  | 29 | 15.2 | II | d 1 c | 1.1 | 1.7 | 144 |
|  |  |  |  | d $2 \mathrm{c}(\mathrm{Z})$ | 2.3 |  |  |
| Dec. | 21 | 14.3 | III | d 1 c | 1.1 | 1.4 | 166 |
|  |  |  |  | d 1.5 c (Z) | 1.7 |  |  |
|  | 25 | 15.4 | III | d 1 c | 1.1 | 1.1 | 170 |
|  |  |  |  | d 1 c ( $Z$ ) | 1.1 |  |  |
|  | 27 | 12.9 | III | d 2 c | 2.3 | 1.7 | 172 |
|  | 28 | 14.7 | II | d 1 c c | 1.1 | 1.7 | 173 |
|  |  |  |  | d 2 c ( Z ) | 2.3 |  |  |
| 84 Jan. | 3 | 14.8 | III | R invis. |  | - | 179 |
|  | 5 | 14.4 | I D | $\begin{align*} & \mathrm{d} 2 \mathrm{c} \\ & \text { d2 } 2 \tag{Z} \end{align*}$ | 2.3 2.3 | 2.3 | 181 |
| May | 28 | 16.0 | I | R invis. |  | $>18$ | 325 |
| June | 14 |  | I | " " |  | > 18 | 342 |
| July | 13 | 16.0 | II | " " |  | " | 371 |
| Aug. | 23 | 15.0 | I | " " |  | " | 412 |
| Sept. | 17 |  |  | " " |  | " | 437 |
| Oct. | 10 | 14.5 | I | " " |  | " | 460 |
| Nov. | 7 | 14.7 | I | " " |  | " | 488 |
| Dec. | 9 | 14.8 | III | " |  | " | 520 |
| 85 June | 9 | 16.1 | I | a 10 e | 18.0 |  |  |
|  | 17 | 15.2 | I | a 10 e? | 18.0 | 18.0 | 710 |
| July | 1 | 16.8 | I | $R$ just vis. |  | $>18$ | 724 |
|  | 13 | 17.3 | I | " " " |  | " | 736 |
| Aug. | 3 | 15.8 | II | R invis. |  | " | 757 |

*) Sky III will explain this.


## METHOD BY STEPS:



| $1800+$ |  |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 88 | June | 30 | 15.8 | II | d 1 R 9 f | 0.6 | 0.9 | 0.8 | 10819 |  |
|  | July | 11 | 16.3 | I | d 3 R 7 f | 2.7 | 2.8 | 2.8 | . 830 |  |
|  |  | 29 | 16.7 | I | d 2 R ! | 2.0 |  | 2.0 | 848 |  |
|  | Aug. | 7 | 15.6 | I | d 3 R 10 f | 1.1 | 2.1 | 1.6 | 857 |  |
|  |  | 12 | 15.6 | I | d 2 R | 2.0 |  | 2.0 | 862 |  |
|  |  | 24 | 15.3 | I | d 4 R | 4.0 |  | 4.0 | 874 |  |
|  |  | 29 | 15.4 | I | d 5 R | 5.0 |  | 5.0 | 879 |  |
|  | Sept. | 6 | 15.0 | , | d 5 R |  |  | 5.0 | 887 |  |
|  | Oct. | 7 | 15.8 | I | for R 4 a c | 9.3 8.8 |  | 9.1 | 918 |  |

7106 S Vulpeculæ

Series IV.
(1900) $19^{\mathrm{h}} 44^{\mathrm{m}} 18^{\mathrm{s}}\left(+2^{\mathrm{s}} .46\right) ;+27^{\circ} 2^{\prime} .3 \quad\left(+0^{\prime} .15\right)$

Periód: $67^{\mathrm{d}} .5$, Periodic inequal.; Variation: $8 \frac{1}{2}^{\mathrm{x}}-9 \frac{1}{2}^{\mathrm{N}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps. | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| d |  | $+26^{\circ} 3679$ | 0.0 | 8.1 BD |
| f |  | $27^{\circ} 3526$ | 7.0 | 9.1 " |
| e |  | $26^{\circ} 3672$ | 8.3 | 9.5 " |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD :


\begin{tabular}{|c|c|c|c|c|c|c|c|c|}
\hline $1800+$ \& Gr. M. T. \& Sky \& Comparisons \& I \& II \& Mean \& $2400000+$ \& Remarks <br>
\hline \multirow[t]{3}{*}{85 July} \& ${ }^{\text {h }}$. ${ }^{\text {a }}$ \& I \& d 6 e \& \& 5.0 \& 5.3 \& 09725 \& <br>
\hline \& \& \& d 8 f \& \& 5.6 \& \& \& <br>
\hline \& 15.5 \& 1 \& d 45 e \& \& 3.3
3.5 \& 3.4 \& 738 \& <br>
\hline \multirow[t]{6}{*}{Aug. $\begin{array}{rr} \\ \\ 1 \\ 1 \\ 1 \\ & 30\end{array}$} \& 15.7 \& I \& d 4 e \& \& 3.3 \& 3.8 \& 758 \& <br>
\hline \& \& \& d 6 f \& \& 4.2 \& \& \& <br>
\hline \& 15.7 \& I \& d 7 e \& \& 5.8 \& 5.7 \& 764 \& <br>
\hline \& 16.1 \& I \& d 78 e \& \& 5.6 \& 5.7 \& 767 \& <br>
\hline \& \& \& d 8 f \& \& 5.6 \& \& \& <br>
\hline \& 15.2 \& I \& d 6 e \& \& 5.0 \& 5.6 \& 784 \& <br>
\hline \multirow[t]{14}{*}{Sept.} \& 15.6 \& I \& d 7 f
d 7 e
d \& \& 4.9
5.8 \& 5.7 \& 786 \& <br>
\hline \& \& \& d 8 f \& \& 5.6 \& \& \& <br>
\hline \& 14.7 \& I \& d 3 e \& \& 2.5 \& 2.7 \& 795 \& <br>
\hline \& 15.6 \& I D \& d ${ }^{\text {d }} 5$ \& \& 2.8
4.2 \& 4.2 \& 799 \& <br>
\hline \& \& \& d 6 f \& \& 4.2 \& \& \& <br>
\hline \& 15.1 \& II D \& d 4 e \& \& 3.3 \& 3.4 \& 802 \& <br>
\hline \& 15.1 \& DD \& d ${ }^{\text {d } 3 \mathrm{f}}$ \& \& 3.5
2.5 \& 2.7 \& 805 \& <br>
\hline \& \& \& d 4 f \& \& 2.8 \& \& \& <br>
\hline \& 15.1 \& DD \& d 3 e \& \& 2.5 \& 2.7 \& 806 \& . <br>
\hline \& 15.0 \& DDD \& d 44 f
d 3 e \& \& 2.8
2.5 \& 3.0 \& 809 \& <br>
\hline \& \& \& d 5 f \& \& 3.5 \& \& \& <br>
\hline \& 14.5 \& " \& d 3 e \& \& 2.5 \& 2.7 \& 810 \& <br>
\hline \& 15.0 \& " III \& d 46
d 6 e \& \& 2.8
5.0 \& 5.0 \& 811 \& <br>
\hline \& \& \& d 7 f \& \& 4.9 \& \& \& <br>
\hline \multirow[t]{18}{*}{Oct.} \& 14.3 \& II \& d 6 e \& \& 5.0 \& 5.0 \& 816 \& <br>
\hline \& \& \& d 7 f \& \& 4.9 \& \& \& <br>
\hline \& 14.4 \& 1 \& d ${ }^{4} 5 \mathrm{e}$ \& \& 3.3 \& 3.4 \& 820 \& not fainter! <br>
\hline \& 14.9 \& I \& d 5 e? \& \& 4.2 \& 4.2 \& 823 \& Eye tired <br>
\hline \& 15.5 \& I \& d 6 f ? ${ }^{\text {d }}$ \& \& 4.2 \& 4.6 \& 824 \& <br>
\hline \& \& \& d 7 f \& \& 4.9 \& \& \& <br>
\hline \& 15.7 \& III \& d 5 e \& \& 4.2 \& 4.2 \& 826 \& <br>
\hline \& 15.2 \& I . \& d 6 f
d 4 e \& \& 4.2
3.3 \& 3.8 \& 828 \& <br>
\hline \& \& \& d 6 f \& \& 4.2 \& \& \& <br>
\hline \& 15.1 \& D \& d 6 e? \& \& 5.0 \& 5.3 \& 830 \& Near ${ }^{\text {d }}$ <br>
\hline \& 15.0 \& D \& d 8 f ? \& \& 5.6
4.2 \& 4.2 \& 831 \& <br>
\hline \& \& ) \& d 6 f ! \& \& 4.2 \& \& \& <br>
\hline \& 14.7 \& DDD \& d 6 e \& \& 5.0 \& 5.3 \& 837 \& <br>
\hline \& 14.1 \& " \& d 8 f
d 6 e

d \& \& 5.6
5.0 \& 5.3 \& 838 \& <br>
\hline \& \& \& d 8 f \& \& 5.6 \& \& \& <br>
\hline \& 14.2 \& III \& d 6 e? \& \& 5.0 \& 5.3 \& 841 \& Fog <br>
\hline \& 14.4 \& I \& d 8 f ? \& \& 5.6
5.0 \& 5.3 \& 843 \& <br>
\hline \& \& \& d 8 f \& \& 5.6 \& \& \& <br>
\hline
\end{tabular}



METHOD BY STEPS:


| 1800+ |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | ---: | :---: | :---: | :---: | :--- | :--- | :--- | :--- | :--- |

(1900) $19^{\mathrm{h}} 46^{\mathrm{m}} 44^{\mathrm{s}}\left(+2^{\mathrm{s}} .31\right) ; \quad+32^{\circ} 39.7 \quad\left(+0^{\prime} .15\right)$

Period: $406^{\mathrm{a}} .02 \pm ; \quad$ Variation: $5^{\mathrm{n}}-13^{\mathrm{n}} .5$
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  | V 10 | $+29^{\circ} 3684$ | 0.0 | (5.3) BD |
| f | III 1 | $+33^{\circ} 3587$ | 4.8 | (5.4) " |
| e |  | $+32^{\circ} 3558$ | 9.2 | (6.5) " |
| d | 3 | +33 ${ }^{\circ} 3602$ | 10.8 | (6.9) " |
| c. | 6 | + $32^{\circ} 3578$ | 16.3 | 8.3 |
| a | 7 | $+32^{\circ} 3589$ | 17.3 | 8.5 |
| b | III 20 | $+32^{\circ} 3583$ | 19.6 | 9.1 |

## Notes:

This star was not on the regular observing list, as it is generally too faint for a 3-inch telescope. The maximu of 1883 was observed in order to obtain some practical knowledge of this variable. The scale of steps could n be derived from the observations, and was therefore based on the steps of the Atlas. The stars $c, a, b$ have the the steps: $13,23,46$, respectively, and the corresponding numbers of the brighter stars were derived from th magnitudes, by means of the formula at the end of the Catalogue. The brightest star $g$ was then chosen for tl zero of the scale, and all the numbers were divided by 10 .

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |


*) In telescope: $\chi>\mathrm{f}$; naked eye: $\chi<\mathrm{f}$.
(1900) $20^{\mathrm{h}} 9^{\mathrm{m}} 30^{\mathrm{s}}\left(+2^{\mathrm{s}} .74\right) ; \quad+16^{\circ} 25^{\prime} .4 \quad\left(+0^{\prime} .18\right)$

Period: $70^{4} .52 \pm$; Variation: $8 \frac{1}{2}^{n}-10^{\mathrm{N}}$.
Comparison Stars:

| Obs. | ASV. | BD . | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| h |  | $+15^{\circ} 4099$ | 0.0 | 8.5 BD. |
| g |  | +16 ${ }^{\circ} 4192$ | 3.0 | 8.7 " |
| f |  | - 4203 | 5.3 | 9.2 " |
| d |  | 4200 | 8.3 | 9.3 " |
| a |  | $+16^{\circ} 4191$ | 13.3 | 9.5 " |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:



| 1800 |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 85 Sept. | 10 | 14.3 | I | d 7 a! |  | 11.8 | 11.8 | 09795 |  |
|  |  |  |  | f 8 a! |  | 11.7 |  |  |  |
|  | 13 | 14.5 | I | d 6 a |  | 11.3 | 11.5 | 798 |  |
|  |  |  |  | f 8 a |  | 11.7 |  |  |  |
|  | 14 | 15.4 | I D | d 7 a |  | 11.8 | 11.8 | 799 |  |
|  | 16 | 15.4 | I D | f 8 a d 7 a |  | 11.7 11.8 | 11.8 | 801 |  |
|  |  |  |  | f 8 a |  | 11.7 |  |  |  |
|  | 17 | 15.1 | II D | d 6 a |  | 11.3 | 11.1 | 802 |  |
|  | 19 | 14.7 | D | f 7 a d 3 a |  | 10.9 9.8 | 9.6 | 804 |  |
|  |  |  | 2 | f 5 a |  | 9.3 | 9.6 | S04 |  |
|  | 20 | 14.8 | DD | d 2 a |  | 9.3 | 8.5 | 805 |  |
|  | 21 | 15.1 | DDD | f 3 a d 2 a |  | 7.7 9.3 | 8.5 | 806 |  |
|  | 21 | 15.1 | D2 | f 3 a |  | 7.7 | 8.5 | 806 |  |
|  | 24 | 14.9 | " | f 4 d ! |  | 6.5 | 6.5 | 809 |  |
|  | 25 | 14.5 | " | $\mathrm{R}=\mathrm{f}$ |  | 5.3 | 4.3 | 810 |  |
|  | 26 | 14.9 | III" | h 8 g h 9 g |  | $(2.4)$ 2.7 | 3.5 | 811 |  |
|  |  |  |  | h 8 f |  | 4.2 |  |  |  |
| Oct. | 1 | 14.1 | II | h 9 f |  | 4.8 | 3.8 | 816 | $\mathrm{f}=\mathrm{g}$ ? |
|  |  |  |  | h 9 g |  | 2.7 |  |  |  |
|  | 5 | 14.1 | I | h 7 f |  | 3.7 2.7 | 3.2 | 820 | $g>f!$ |
|  | " | 15.7 | " | h 8 ¢ |  | 4.2 | 3.5 |  |  |
|  |  |  |  |  |  | 2.7 |  |  |  |
|  | 8 | 14.8 | I | h 9 g ? |  | 2.7 | 3.4 | 823 | Eye tired. |
|  | 9 | 15.3 | I | g 9 f |  | 5.1 | 4.9 | 824 |  |
|  |  |  |  | g 3 d |  | 4.6 |  |  |  |
|  | 10 | 15.1 | III | g 9 f |  | 5.1 | 5.1 | 825 |  |
|  | 11 | 15.3 | III | g 4 d f 1 d ! |  | 5.1 | 5.9 | 826 |  |
|  |  |  |  | g 6 d ! |  | 6.2 |  |  |  |
|  | 13 | 15.1 | I | g 5 d ! |  | 5.6 | 5.8 | 828 |  |
|  |  |  |  | f 2 d ! |  | 5.9 |  |  |  |
|  | 15 | 15.0 | D | g 6 d |  | 6.2 | 6.4 | 830 | Near D |
|  | 16 | 14.9 |  | f 4 d f 3 d |  | 6.5 6.2 | 6.2 | 831 |  |
|  | 22 | 14.6 | DDD | d 1 a? |  | 8.8 | (8.3) | 837 |  |
|  |  |  |  | f 3 a |  | 7.7 |  |  |  |
|  | 23 | 14.1 | " | f 9 d |  | 7.5 | 7.5 | 838 |  |
|  | 26 | 14.3 | III | f 7 d ? |  | 7.4 | 7.6 | 841 |  |
|  |  |  |  | f 3 a? |  | 7.7 |  |  |  |
|  | 28 | 14.4 | I | f 7 d |  | 7.4 | 7.3 | 843 |  |
|  | 29 | 16.1 |  | g 8 d f 10 d |  | 7.2 8.3 | 8.0 | 844 |  |
| Nov. |  |  |  | f 3 a |  | 7.7 |  |  |  |
|  | 4 | 15.2 | II | f 9 d ! |  | 7.5 | 7.6 | 850 |  |
|  |  |  |  | f 3 a |  | 7.7 |  |  |  |
|  | 9 | 15.2 | II | $f 8 \mathrm{~d}$ |  | 7.7 | 7.4 | 855 |  |
|  |  |  |  | g 9 d |  | 7.8 |  |  |  |
|  |  |  |  | g 3 a |  | (6.1) |  |  |  |



METHOD BY STEPS:

| 87 May | 24 | 17.0 |  | f 1 R 2 d g 2 R 2 d | 6.3 5.7 | 6.3 5.7 | 6.0 | 416 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| June | 15 | 16.3 | I | ${ }_{\text {d }} 1 \mathrm{R} 3 \mathrm{R}$ | 5.7 9.8 | 9.7 | 9.7 | 438 |  |
|  | 19 | 16.0 | I | h 1 R 1 g | 1.5 | 1.5 | 2.0 | 442 |  |
|  |  |  |  | h 1 R 2 f | 2.2 | 1.8 |  |  |  |
|  | 21 | 15.8 |  | R4d ${ }_{\text {h } 2 \mathrm{R}} 1 \mathrm{~g}$ | $(4.3)$ 2.0 | 2.0 | 2.4 | 444 |  |
|  |  |  |  | R2f | 3.3 |  | 2.4 | 444 |  |
| July | 10 | 15.8 | I | f 2 R 1 d | 7.3 | 7.3 | 7.3 | 463 |  |
|  | 15 | 15.9 | I | f1 R 2 d | 6.3 | 6.3 | 7.1 | 468 |  |
|  |  |  |  | f1R3a | 8.3 | 7.3 |  |  |  |
|  | 22 | 15.9 | I | h2R0f | 3.6 |  | 2.8 | 475 |  |
|  |  |  |  | $\mathrm{R} 1 . \mathrm{g}$ | 2.0 |  |  |  |  |
| Aug. | 6 | 15.3 | II | f $2 \mathrm{R} 1 \mathrm{C}^{\text {f }} \mathrm{R} 4 \mathrm{a}$ | 7.3 8.3 | 7.3 8.0 | 7.7 | 490 |  |
| Sept. | 6 | 16.2 | II | g 1 R 2 f | 3.7 | 3.8 | 4.1 | 521 |  |
|  |  |  |  | g 1R3d | 4.7 | 4.3 |  |  |  |
|  | 15 | 15.5 | I |  | 5.1 5.3 | 4.8 | 5.1 | 530 |  |
|  | 18 | 15.4 | I | g 1 R 2 d | 5.1 | 4.8 | 5.1 | 533 |  |
| Oct. | 12 | 14.8 |  | R 0 f f 2 R 1 d | 5.3 7.3 | 7.3 | 6.8 | 557 |  |
|  |  |  |  | g 2 R 1 d | 6.2 | 6.5 |  |  |  |
|  | 18 | 14.8 | II | $\mathrm{f}^{4} \mathrm{R} 5 \mathrm{a}$ | 8.8 | 8.9 | 9.0 | 563 |  |
|  |  |  |  | d 1 R g 2.5 R 1.5 f | 9.3 |  |  |  |  |
| Nov. | $\begin{aligned} & 11 \\ & 17 \end{aligned}$ | $\begin{aligned} & 15.5 \\ & 15.0 \end{aligned}$ | $\mathrm{I}_{\text {I }}$ | ${ }_{\text {g } 2.5 \mathrm{~F}}^{\mathrm{h}} \mathrm{l} 1.5 \mathrm{f}$ | 4.6 |  | 4.5 3.5 | 587 593 |  |
|  |  |  |  | h 3 R 1 g h 3 l f | 2.5 3.7 | 2.2 4.0 | 3.5 | 593 |  |
|  |  |  |  | h 3 R 3 d | 4.2 | 4.2 |  |  |  |
| 88 May | 28 | 16.3 | I | d 3 R 0 a | 12.3 |  | 12.3 | 786 |  |
| June | 2 30 | 16.0 15.9 | I | d 1 R 2 a f 2 d | 10.3 7.3 | 10.0 7.3 | 10.2 7.7 | 791 819 |  |
|  | 30 | 15.9 | II |  | 7.3 8.3 | 7.3 8.0 | 7.7 | 819 |  |
| July | 11 | 16.5 | I | g 1 R 1 f | 4.1 | 4.1 | 4.0 | 830 |  |
|  |  |  |  | h3R1f | 3.7 | 4.0 |  |  |  |
|  | 29 | 16.8 | I | f 1 R 2 d | 6.3 | 6.3 | 6.3 | 848 | Seeing poor |


|  | 1800 |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 88 | Aug. | 7 | 15.8 | I | d 4 R 1 a! | 12.3 | 12.3 | 12.3 | $10 \cdot 857$ |  |
|  |  | 12 | 15.9 | I | a 1 R | 14.3 |  | (14.3) | 862 |  |
|  |  | 24 | 15.4 | I | f1R1d | 6.8 | 6.8 | 6.8 | 874 |  |
|  |  | 29 | 15.6 | I | g 2 R 1 f | 4.5 | 4.5 | 4.4 | 879 |  |
|  | Sept. | 6 | 15.1 | I | R 4 d f 1 R 3 d | 4.3 5.8 | 6.0 | 5.9 | 887 |  |
|  | Oct. | 7 | 16.0 | I | d 1 R 4 a | 9.3 | 9.3 | 9.3 | 918 |  |

(1900) $20^{\mathrm{h}} 10^{\mathrm{m}} 5^{\mathrm{s}}\left(+2^{\mathrm{s} .90)} ; \quad+8^{\circ} 47^{\prime} .1 \quad\left(+0^{\prime} .18\right)\right.$

Period: $285^{\mathrm{a}} .5$, periodic inequal.?; Variation: $8^{\mathrm{n}}-12^{\mathrm{n}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| e | 1 | $+8^{\circ} 4393$ | 0.0 | (6.7) |
| d | 2 | $9^{\circ} 4452$ | 7.0 | 8.4 |
| c | 4 | $8^{\circ} 4383$ | 10.5 | 8.7 |
| a | 7 | 4389 | 14.8 | 9.0 |
| b | 10 | 4385 | 16.8 | 9.2 |
| f | 15 | 4384 | 20.8 | 9.8 |
| h | 14 | $+8^{\circ} 4388$ | (20.8) | 9.6 |
| g | 27 | ........ | 23.8 | 11.0 |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:



| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |

METHOD BY STEPS :

| 87 May June | 24 | 17.2 |  | R barely vis. |  |  | $\geq 24$ | 10416 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 15 | 16.5 | I | a 1 R 1 b | 15.8 | 15.8 | 15.8 | 438 |
|  | 19 | 16.1 | I | c 3 R 2 b | 14.2 | 14.3 | 13.5 | 442 |
|  |  |  |  | c 3 R 3 a | 12.7 | 12.7 |  |  |
|  | 21 | 15.8 |  | c 2 R 3 b | 13.1 | 13.0 | 13.0 | 444 |
|  |  |  |  | R 2 a | 12.8 |  |  |  |
|  | 23 | 15.8 | I | d 3 R 3 a | 10.4 | 10.9 | 10.6 | 446 |
|  |  |  |  | R 0 c | 10.5 |  |  |  |
|  | 25 | 15.6 | II | d 3 R 3 a | 10.4 | 10.9 | 10.3 | 448 |
|  |  |  |  | R 1 c | 9.5 |  |  |  |
| July | 10 | 15.9 | I | R 2 d | 5.0 |  | 6.3 | 463 |
|  |  |  |  | R 3 c | 7.5 |  |  |  |
|  | 15 | 16.0 | I | e 5 R 2 d | 5.0 | 5.0 | 5.7 | 468 |
|  |  |  |  | e 5 R 3 c | 6.3 | 6.6 |  |  |
|  | 22 | 16.0 | I | e 4 R 3 d | 4.0 | 4.0 | 4.7 | 475 |
|  |  |  |  | e 4 R 4 c | 5.3 | 5.3 |  |  |
| Aug. | 6 | 15.5 | I | R 1 d | 6.0 |  | 6.3 | 490 |
|  |  |  |  | R 4 c | 6.5 |  |  |  |
| Sept. | 6 | 16.3 | II | b 2 R 2 f | 18.8 | 18.8 | 18.8 | 521 |
|  | 16 | 15.1 | II | f1R2g | 21.8 | 21.8 | 21.8 | 531 |
|  | 18 | 15.5 | I |  | 21.8 | 21.8 | 21.8 | 533 |
| Oct. | 12 | 14.9 |  | $R$ barely vis. |  |  | $>24$ | 557 |
| 88 June | 2 | 16.1 | I | a 1 R 1 b | 15.8 |  | 15.4 | 791 |
|  |  |  |  | c 3 R 1 b | 14.7 | 15.2 |  |  |

Nova Andromedæ 1885

$$
\begin{gathered}
(1900) 0^{\mathrm{h}} 37^{\mathrm{m}} 15^{\mathrm{s}}\left(+3^{\mathrm{s}} .26\right) ;+40^{\circ} 43^{\prime} .2 \quad\left(+0^{\prime} .33\right) \\
\text { Variation : } 7^{\mathrm{m}}-<13^{\mathrm{M}} \\
\text { Comparison Stars: }
\end{gathered}
$$

| Obs. | ASV. | BD . | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| A | ..... | $+40^{\circ} 158$ | ......... | 7.5 BD. |
| D | ...... | 151 | ......... | 8.9 |
| B | ...... | 156 | ......... | 9.0 |
| C | ...... | $+40^{\circ} 154$ | ......... | 9.0 |

## Notes:

The 3-inch equatorial was evidently unable to show the variations of this star so as to give a knowledge of its light curve. The feeble attempt, which the novelty of this phenomenon seemed to demand, is here reproduced, with a reduction to the magnitudes of the BD . scale.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:



## II. Observations made by Arthur Zaiser from 1884 to 1887.

These observations were made by Arthur Zaiser, a student of the College in Prairie du Chien, Wisconsin. They were all made with the naked eye, aided sometimes by an opera-glass. The value of one step differs widely for the different variables, especially for the brighter ones, and amounts on the average to almost 0.3 of a magnitude.

1411
$\lambda$ Tauri
Series V.
(1900) $3^{\mathrm{n}} 55^{\mathrm{m}} 8^{\mathrm{s}}\left(+3^{\mathrm{s}} .32\right) ;+12^{\circ} 12^{\prime} .5 \quad\left(+0^{\prime} .17\right)$

Period: $3^{\mathrm{d}} 22^{\mathrm{h}} 52^{\mathrm{m}} .02$; Var.: $3^{\mathrm{x}} .4-4^{\mathrm{n}} .2$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
|  | 1 | $+8^{\circ} 511$ | 0.0 | 3.8 |
| $\sigma$ | 27 | $15^{\circ} 612$ | 0.0 | 3.8 |
| $\gamma^{\circ}$ | 2 | $9^{\circ} 439$ | 3.8 | 3.8 |
| $\mu$ | 26 | $8^{\circ} 657$ | 4.5 | 4.3 |
| f | 3 | $+12^{\circ} 486$ | 10.0 | 4.3 |

Notes:
The observations of this star, which is of the Algol-type, were made for practice rather than with a view of obtaining exact times of minima. For this reason no particular care was taken of applying a correction to the watch, which may have been two or three minutes wrong.

| $1800+$ | Gr. M.T. | Sky. | Comparisons | I | .II | Mean | 2400000 |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |$\quad$ Remarks

DECIMAL METHOD :

| 84 Oct. | 11 | $15^{\mathrm{n}} 16^{\mathrm{m}}$ | II | 03 f | 3.0 | 2.6 | 09461 | F G. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 12 | 1520 | I | - 6 | 2.3 3.0 | 2.9 | $462$ |  |
|  | 12 | 1520 | 1 | 063 $07 \xi$ | 3.0 2.7 | 2.9 | 462 |  |
|  | 16 | 1538 | II | 03 f | 3.0 | 2.9 | 466 |  |
|  |  |  |  | $07 \%$ | 2.7 |  |  |  |
| Nov. | 7 | 1544 | I | 03 f | 3.0 | 2.9 | 488 |  |
|  |  |  |  | 078 | 2.7 |  |  |  |
|  | 18 | 1512 | I | o 3 f | 3.0 | 3.0 | 499 |  |
|  |  |  |  | - 5 § | 3.0 |  |  |  |
| Dec. | 9 | 1422 | II | o 2 f | 2.0 | 2.4 | 520 |  |
|  |  |  |  | o 7 § | 2.7 |  |  |  |
|  | 18 | $15 \quad 5$ | I | o 3 f | 3.0 | 3.0 | 529 | $4^{\text {n }}$ after Eph. Min. |
|  |  |  |  | o 8 今 | 3.0 |  |  |  |
| 85 Jan. | 9 | 1535 | I | o 2 f | 2.0 | 2.7 | 551 |  |
|  |  |  |  | - 9 \% | 3.4 |  |  |  |
|  | 11 | 1515 | I | o 3 f | 3.0 | 2.7 | 584 | $4^{\text {b }}$ before Eph. Min. |
|  |  |  |  | - 6 \% | 2.3 |  |  |  |
|  | 7 | 1442 | I | o 4 f | 4.0 | 3.7 | 608 | F G. $2^{\text {n }}$ after Eph. Min |
|  |  |  |  | $\bigcirc 9 \xi$ | 3.4 |  |  |  |
|  | 9 | $15 \quad 9$ | II | - 2 f | 2.0 | 2.0 | 610 | " |
|  |  |  |  | $05 \xi$ | 1.9 |  |  |  |
| Oct. | 8 | 1542 | I | 03 f | 3.0 | 3.7 | 823 | . |
|  |  |  |  | $\xi 1 \mathrm{f}$ | 4.4 |  |  |  |
|  | 9 | 1546 | I | o 2 f | 2.0 | 2.9 | 824 |  |
|  |  |  |  | $\xi 0 \mathrm{f}$ | 3.8 |  |  |  |
| Nov. | 9 | 1513 | I | o 3 f | 3.0 | 3.7 | 855 |  |
|  |  |  |  | $\xi 1 \mathrm{f}$ | 4.4 |  |  |  |
| Dec. | 6 | 1516 | I | - 2 f | 2.0 | 2.9 | 882 |  |
|  |  |  |  | $\xi 0 \mathrm{f}$ | 3.8 |  |  |  |
|  | 11 | 1516 | I | - 2 f | 2.0 | 2.0 | 887 |  |
|  |  |  |  | - 5 \% | 1.9 |  |  |  |
| S6 Jan. | 6 | 1333 | III | 01 f | 1.0 | 1.5 | 913 |  |
|  |  |  |  | $05 \%$ | 1.9 |  |  |  |
|  | 9 | 1510 | I | 07 f $\xi 5 \mathrm{f}$ | 7.0 6.9 | 7.0 | 916 | $5 \frac{1}{2}^{\text {b }}$ before Eph. Min. |
| Febr. | 1 | 1510 | I | - 2 f | 2.0 | 2.0 | 939 |  |
|  |  |  |  | 0 $5 \xi$ | 1.9 |  |  |  |
|  | 2 | 1342 | II | o 9 f | 9.0 | 8.6 | 940 | Eph. Hel. Min. $14^{\text {h }} 7^{\text {m }}$ |
|  |  |  |  | $\xi 7 \mathrm{f}$ | 8.1 |  |  |  |
|  | " | $14 \quad 9$ | " | - 9.5 f | 9.5 | 9.2 | " |  |
|  |  |  |  | $\xi 8 \mathrm{f}$ | 8.8 |  |  |  |
|  | 8 | 1455 | III | o 3 f | 3.0 | 3.4 | 946 |  |
|  |  |  |  | $\xi 0 \mathrm{f}$ | 3.8 |  |  |  |
|  | 22 | 146 | II | o 4 f | 4.0 | 4.2 | 960 | F G. |
|  |  |  |  | $\xi 1 \mathrm{f}$ | 4.4 |  |  |  |
| Mar. | 2 | 1528 | I | o 2 f | 2.0 | 2.2 | 968 |  |
|  |  |  |  | $06 \xi$ | 2.3 |  |  |  |


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | 2400000 | Remarks |
| :--- | ---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 86 Oct. 29 | $15^{\mathrm{h}} 15^{\mathrm{ma}}$ | I | 03 f  <br> 0 7. |  | 3.0 <br> 2.7 | 2.9 | 10209 |  |


| 87 Mar. 24 | 1525 | II | $\gamma 2 \lambda 2 \mu$ | 2.3 | 2.3 | 2.3 | 355 | F G. Eph. Hel. Min. $15^{\mathrm{h}} 43^{\mathrm{m}}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 1538 | " | r3ג2 $\mu$ | 2.7 | 2.8 | 2.8 |  |  |
| " | 1549 | " | r3^1 $\mu$ | 3.4 | 3.3 | 3.4 | " | " |
| " | 1558 | " | $\gamma 4 \lambda 0.5 \mu$ | 4.0 | 4.0 | 4.0 | " | " |
| " | 169 | " | $\gamma 4 \lambda 0.5 \mu$ | 4.0 | 4.0 | 4.0 | " |  |
| " | 1625 | " | $\gamma 5^{2} \lambda 0.5 \mu$ | 4.1 | 4.5 | 4.3 | " | " Low. |
| 28 | 1433 | II | $\gamma 4 \lambda 0 \mu$ | 4.3 | 4.5 | 4.4 | 359 | F G. Eph. Hel. Min. $14^{\mathrm{h}} 35^{\mathrm{m}}$ |
| " | 1439 | " | $\gamma 4 \lambda 0.5 \mu$ | 4.0 | 4.0 | 4.0 |  | " |
| " | 1451 | " | ¢ 420 \% | 4.3 | 4.5 | 4.4 | " | " |
| " | 1510 | " | $\gamma 4 \lambda 1 \mu$ | 3.6 | 3.8 | 3.7 | " | " |
| " | 1525 | " | $\gamma 3 \lambda 1.5 \mu$ | 3.0 | 3.3 | 3.2 | " | " |
| " | 1535 | " | r3 $22 \mu$ | 2.7 | 2.8 | 2.8 | " | " |

$2098 \quad \alpha$ Orionis $\quad$ Series V.

Irregularly periodic; Variation: $1^{n}-1^{x} .4$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| $\beta$ | 39 | - $8^{\circ} 1063$ | 2.4 | 0.3 |
| ${ }^{\circ}$ | 9 | +16 ${ }^{\circ} 629$ | 3.4 | 1.1 |
| $r$ | 43 | + $6^{\circ} 919$ | 6.2 | 1.6 |

Notes:
The comparison star $a$ is a Tauri. Five comparisons with Sirius were omitted, as this star is too bright and too different in color to give useful results. The observations may not contribute much to the knowledge of the irregular variations of this star, since it is not a suitable object for the method by steps, and seems to require a photometer.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:


| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Remarks |  |  |  |  |  |  |  |

METHOD BY STEPS:

| Jan. | 28 | 14.8 | II | $\beta .0 \times 2 \gamma$ | 3.3 |  | 3.3 | 10300 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Febr. | 12 | 14.4 | I | B1 $\alpha 4 \gamma$ | 2.8 | 3.2 | 3.0 | 315 |
|  | 16 | 15.1 | II | $\beta 0 \sim 5 \gamma$ | 1.8 |  | 1.8 | 319 |
| Mar. | 13 | 14.7 | II | B $0 \times 6 \gamma$ | 1.3 |  | 1.3 | 344 |
|  | 17 | 14.4 | II | $\beta 0.50 .4 \gamma$ | 2.6 | 2.8 | 2.7 | 348 |
|  | 24 | 15.7 | II | ${ }_{\alpha} 1 \beta$ | 1.4 |  | 1.8 | 355 |
|  | 28 | 15.2 | II | a $4 \gamma$ ¢ 0 $a_{a} 3 \gamma$ | 2.2 2.8 |  | 2.8 | 359 |


| 2509 | $\zeta$ Geminorum | Series V |
| :--- | :--- | :--- |
| $(1900) 6^{\mathrm{n}} 58^{\mathrm{m}} 11^{\mathrm{s}}\left(+3^{\mathrm{s}} .56\right) ;+20^{\circ} 43^{\prime} .0$ | $\left(-0^{\prime} .09\right)$ |  |

Period: $10 .{ }^{d} 15382 ; \quad$ Variation: $3^{\mathrm{n}} .7-4^{\mathrm{x}} .5$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
| $\lambda$ | 74 | $+16^{\circ} 1443$ | 0.0 | 3.7 |
| $\delta$ | 75 | $22^{\circ} 1645$ | 2.1 | 3.5 |
| d | 69 | $+21^{\circ} 1405$ | 15.0 | 5.3 |

Notes:
The step-interval between $\delta$ and $d$ is too large to yield concordant results. The table of comparison stars in the Atlas (Ser. V, ch. VI) shows that the comparison star $d$ has not been employed by any other observer.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD :



## METHOD BY STEPS:



## Series V.

(1900) $14^{\mathrm{b}} 55^{\mathrm{m}} 38^{\mathrm{s}}\left(+3^{\mathrm{a}} .20\right) ;-8^{\circ} 7^{\prime} .3\left(-0^{\prime} .24\right)$

Period: $2^{\mathrm{a}} 7^{\mathrm{h}} 51^{\mathrm{m}} 22^{\mathrm{a}} .8$; Variation: $5^{\mathrm{m}} .0-6^{\mathrm{M}} .2$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| a | 6 | $-3^{\circ} 3696$ | 0.0 | 4.6 |
| b | -9 | $-1^{\circ} 2991$ | 1.9 | 5.0 |
| c | 9 | $-4^{\circ} 3783$ | 5.4 | 6.0 |

Notes:
Although the observations of this Algol-star were not made with a systematic plan to obtain the exact time of the minima, still they determine the ascending branch of the light curve on three different occasions.

| $1800+$ | Gr. M. I.. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:

| 84 Sept. | 11 17 | $14^{\mathrm{h}} 14^{\mathrm{m}}$ 1348 | I I | a 1 c a 4 b a 9 c b 5 c | 0.5 0.2 4.9 3.7 | 0.4 4.3 | 09431 437 | F G. used throughout. $2^{\text {n }}$ after Eph. Min. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 85 May | 5 | 1545 | II | a 1 c | 0.5 | 0.8 | 667 |  |
|  | 8 | 162 | I | a 5 b a 3 c | 1.0 1.6 | 2.0 | 670 |  |
|  |  |  |  | b 1 c | 2.3 | 2.0 | 67 |  |
|  | 13 | 16.46 | III | a 4 c | 2.2 | 2.2 | 675 | *) |
|  | 4 | 1540 |  | b 1 c | 2. 2 | 2.0 |  |  |
| June | 4 | 1540 | 1 | a 3 c b 1 c | 1.6 2.3 | 2.0 | 697 |  |
|  | 8 | 1556 | I | a 2 c b 0 c | 1.1 1.9 | 1.5 | 09701 |  |
| 86 Apr. | 27 | 1527 | II | a 1 c | 0.5 | 0.9 | 10024 |  |
| May | 1. | 1451 | I | a 7 b a 8 c | 1.3 4.3 | 4.0 |  |  |
|  | 1. |  | I | a <br> b 5 c | 4.3 3.7 | 4.0 | 028 |  |
|  | " | 150 | " | a 8.5 c | 4.6 | 4.2 | " | Eph. Hel. Min. $15^{\text {b }} 0^{\text {m }}$ |
|  | " | 1513 | " | b 5 c a 8 c | 3.7 4.3 | 4.0 | " |  |
|  |  |  |  | b 5 c | 3.7 |  |  |  |
|  | " | 1533 | " | a 8 c | 4.3 | 4.0 | " |  |
|  |  |  |  | b 5 c | 3.7 |  |  |  |
|  | " | - 1558 | " | a 7 c | 3.8 | 3.8 | " |  |
|  |  |  |  | b 6 c | 4.0 |  |  |  |
|  | " | 167 | " | a 6 c | 3.2 | 3.5 | " |  |
|  |  |  |  | b 5 c | 3.7 |  |  |  |
|  | " | 1617 | " | a 5 c b 4 c | 2.7 3.3 | 3.0 | " |  |
|  | 6 | 158 | III | a 2 c | 1.1 | 1.1 | 033 |  |
|  |  |  |  | a 5 b | 1.0 |  |  |  |
|  | 18 | $15 \quad 5$ | I D | a 3 c | 1.6 | 1.6 | 045 | Near ( |
|  |  |  |  | a 8 b | 1.5 |  |  |  |
|  | 27 | 1534 | I | a 1 c | 0.5 | 0.8 | 054 |  |
|  | 29 | 161 | III | a 6 b a 3 c | 1.1 | 1.8 | 056 |  |
|  |  |  |  | a 10 b | 1.9 |  |  | $3{ }^{\text {a }}$ after Eph. Min. |
| June | 2 | 1543 | I | a 1 c | 0.5 | 0.8 | 060 |  |
|  |  |  |  | a 6 b | 1.1 |  |  |  |
|  | 17 | 1515 | I DD | a 3 c | 1.6 | 2.0 | 075 |  |
|  |  |  |  | b 1 c | 2.3 |  |  |  |

## METHOD BY STEPS:


*) The original has : a $1 \mathrm{c}, \mathrm{b} 4 \mathrm{c}$, which must be an error in recording.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 87 May | $15^{\mathrm{m}} 5^{\text {mi }}$ | II | a $4 \delta 1 \mathrm{c}$ | 4.2 | 4.3 | 4.2 | 10412 |  |
|  | 1513 |  | b 2 o 1 c | 4.2 | 4.2 |  |  |  |
|  | 1513 |  |  | 4.2 4.2 | 4.3 4.2 | 4.2 |  | Ephem. Hel. Min. $15^{\mathrm{h}} 10^{\mathrm{ma}}$ |
|  | 1524 | " | a 3 or 1 c | 3.7 | 4.1 | 4.1 | " |  |
|  |  |  | b 2 or 1 c | 4.2 | 4.2 |  |  |  |
|  | 1537 | " | a 3 \% 2 c | 3.2 | 3.2 | 3.2 | " |  |
|  | 1549 | " | a 2 o 3 c | 2.2 | 2.2 | 2.5 | " |  |
|  |  |  | b $1 \delta 3 \mathrm{c}$ | 2.7 | 2.8 |  |  |  |
|  | 1459 | III D | a 5 \% 1.5 c | 4.5 | 4.1 | 4.1 | 419 | Ephem. Hel. Min. $14^{\text {n }} 43^{\text {an }}$ |
|  | 158 | " | b 301.0 c $\mathrm{a} 4 \delta 2 \mathrm{c}$ | 3.5 3.7 | 4.2 | 3.4 | " |  |
|  |  |  | b 1 ¢ 2 c | 3.2 | 3.1 |  |  |  |
|  | 1519 | " | a $4 \delta 2 \mathrm{c}$ | 3.7 | 3.6 | 3.7 | " |  |
|  | 1529 | " | b 2 o 2 c | 3.7 | 3.7 |  |  |  |
|  |  |  | b $2 \delta 2.5 \mathrm{c}$ | 3.4 | 3.4 |  |  |  |
|  | 1550 | " | a 2.5 o 3 c | 2.5 | 2.5 | 2.6 | " |  |
|  |  |  | blo 3 c | 2.7 | 2.8 |  |  |  |

$6181 \quad \alpha$ Herculis Series V.
(1900) $17^{\mathrm{h}} 10^{\mathrm{m}} 5^{\mathrm{s}}\left(+2^{\mathrm{s} .73)} ; \quad+14^{\circ} 30^{\prime} .2\left(-0^{\prime} .07\right)\right.$

Irregular; Variation: $3^{n} .1-3^{\mathrm{n}} .9$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :--- | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| a | 69 | $+12^{\circ} 3252$ | 0.0 | 2.1 |
| $\delta$ | 52 | $+25^{\circ} 3221$ | 1.4 | 3.1 |
| $\kappa$ | 68 | $+9^{\circ} 3298$ | 2.5 | 3.5 |

## Notes:

Comparison star $a$ is $\alpha$ Ophinchi, and $\kappa$ is $\kappa$ Ophinchi, while $\delta$ belongs to the constellation Hercules. Another comparison star, fainter than $\kappa$, should have been chosen, at least for the observations in 1887, for which formula II cannot now be employed. Charta X . of the Atlas will show that $\gamma, \varepsilon, \xi$ Herculis have been used by other observers.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD :



METHOD BY STEPS:



6202 u Herculis Series V.
(1900) $17^{\mathrm{h}} 13^{\mathrm{m}} 38^{\mathrm{s}} \quad\left(+2^{\mathrm{s}} .21\right) ;+33^{\circ} 12^{\prime} .3 \quad\left(-0^{\prime} .07\right)$

Irregularly periodic; Variation: $4^{\mathrm{x}} .6-5^{\mathrm{x}} .4$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| ---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| $\varepsilon$ | 48 | $+31^{2} 2947$ | $\ldots \ldots$. | 3.8 |
| w | 56 | $+32^{2} 2896$ | $\ldots \ldots$. | 5.3 |
| c | 58 | $+34^{\circ} 2971$ | $\ldots \ldots \ldots$ | 5.8 |

## Notes:

The observations, being all made by the decimal method, afford no means of establishing an independent scale. They were reduced directly to the magnitudes of the HP. scale, which is given in the last column of the abvoe table.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:


$6758 \quad \beta$ Lyræ Series V.
$(1900) 18^{\mathrm{h}} 46^{\mathrm{m}} 23^{\mathrm{s}}\left(+2^{\mathrm{n}} .21\right) ;+33^{\circ} 14^{\prime} .8 \quad(+0.07)$
Period: $12^{\mathrm{d}} 21^{\mathrm{n}} 47^{\mathrm{m}} 23^{\mathrm{s}} .72+$; Variation: $3^{\mathrm{N}} .4-4^{\mathrm{n}} .5$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| $\gamma$ | 23 | $+32^{\circ} 3286$ | 0.0 | 3.3 |
| ¢ | $\left\{\begin{array}{l}20 \\ \hline 1\end{array}\right.$ | $\left.+36^{\circ} 3307\right\}$ | 3.0 | $\{\quad 5.6$ |
| , | $\left\{\begin{array}{l}21 \\ 17\end{array}\right.$ | +36 +3319 | 3.0 | 1 4.5 |
| \% | 17 | +3703223. | 4.3 | 4.2 \& 5.8 |

## Notes:

The last two columns of the above table show the advantage of establishing a scale for the comparison star: from the observations themselves, independently of photometric measures made by other observers and by other means. A computation of the combined effect of two component stars would bring a new element of uncertainty into the direct estimates of the observer.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |




METHOD BY STEPS:
87 Apr. 18


| 0.9 | 1.1 |
| :--- | :--- |
|  |  |

$380 \mid$ F G.


7124
$\eta$ Aquilæ
Series V.
(1900) $19^{\mathrm{n}} 47^{\mathrm{m}} 23^{\mathrm{s}}\left(+3^{\mathrm{s}} .06\right) ; \quad+0^{\circ} 44^{\prime} .9 \quad\left(+0^{\prime} .15\right)$

Period: $7^{\mathrm{d}} .176381$; Variation: $3^{\mathrm{x}} .5-4^{\mathrm{x}} .7$
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| $\gamma$ | 63 | $+10^{\circ} 4043$ | 0.0 | 2.8 |
| $\beta$ | 70 | $+6^{\circ} 4357$ | 3.0 | 3.8 |
| $\mu$ | 52 | $+7^{\circ} 4132$ | 5.8 | 4.5 |

## Notes:

Since the light curve of this star is well determined, these observations, although few in number, may be of use in supplementing other series.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

## DECIMAL METHOD:



## METHOD BY STEPS:


(1900) $21^{\mathrm{h}} 40^{\mathrm{m}} 27^{\mathrm{s}}\left(+1^{\mathrm{s}} .83\right) ;+58^{\circ} 19^{\prime} .3 \quad\left(+0^{\prime} .27\right)$

Irregularly periodic; Variation: $4^{\mathrm{x}}$ ? $-6^{\mathrm{n}}$ ?
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magı. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| $\varsigma$ | 36 | $+57^{\circ} 2475$ | 3.54 | 3.7 |
| $\iota$ | 40 | $+65^{\circ} 1814$ | 3.62 | 3.7 |
| $\varepsilon$ | 38 | $+56^{\circ} 2741$ | 4.24 | 4.2 |

## Notes:

The scale of steps is in this case not deduced from the observations, but is simply the scale of the H. P magnitudes (vol. XIV). The reductions were made in 1889, in order to compare these observations with the simultaneous ones of Mr. Gore, published in the Proceedings of the Royal Irish Academy, 3d Ser., vol. I, No. I As the latter were reduced to the H. P. scale, the former were reduced to the same. The observations made by the method of steps show, however, that they do not fit well into this scale.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | Remarks

DECIMAL METHOD:




METHOD BY STEPS :

| Jan. | 28 | 12.3 |  | \% $4 \mu 68$ | 2.89 | 3.82 | 3.36 | 300 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Feb. | 12 | 14.9 | I | $53 \mu 78$ | 1.89 | 3.75 | 2.82 | 315 |  |
|  | 16 | 15.3 | 'II | $\zeta 2 \mu 5=$ | 2.39 | 3.74 | 3.07 | 319 |  |
| Mar. | 13 | 14.9 | II |  | 4.89 | 4.04 | 4.47 | 344 |  |
|  | 17 | 14.6 | I | ${ }_{6} 1 \mu 38$ | 2.89 | 3.71 | 3.30 | 348 |  |
|  | 24 | 15.9 | II | $\zeta 1 \mu 3=$ | 2.89 | 3.71 | 3.30 | 355 | F G |
|  | 28 | 15.3 | II | \% $1 \mu 4$ \% | 2.39 | 3.68 | 3.04 | 359 |  |
| Apr. | 18 | 15.6 | II | $\zeta 1 \mu 38$ | 2.89 | 3.71 | 3.30 | 380 | " |
|  | 20 | 15.3 | I | \% $2 \mu 2$ \% | 3.89 | 3.89 | 3.89 | 382 |  |
|  | 25 | 16.0 | I | $\zeta 1 \mu 3 \varepsilon$ | 2.89 | 3.71 | 3.30 | 387 | " |
| May | 15 | 15.8 | II | $\zeta^{\prime} 1 \mu 1 \varepsilon$ | 3.89 | 3.89 | 3.89 | 407 | " |
|  | 19 | 15.9 | II | $\zeta 1 \mu 3$ в | 2.89 | 3.71 | 3.30 | 411 |  |
| June | 14 | 15.9 | II | ${ }_{51 \mu} 1^{\prime}{ }_{8}$ | 2.89 | 3.71 | 3.30 | 437 |  |
|  | 17 | 16.5 | I | ${ }_{\zeta} 1 \mu 28$ | 3.39 | 3.77 | 3.58 | 440 |  |
|  | 23 | 15.6 | I | $\zeta 2^{\mu} 2$ 2 | 3.89 | 3.71 | 3.80 | 446 |  |
| July | 6 | 16.7 | D | $\zeta 2 \mu 2.5$ ع | 3.64 | 3.94 | 3.79 | 459 |  |
|  | 10 | 16.2 | I | $\zeta 2 \mu 3 \varepsilon$ | 3.39 | 3.84 | 3.62 | 463 |  |
| Aug. | 15 | 16.8 | II | $\zeta 2 \mu 2 \varepsilon$ | 3.89 | 3.89 | 3.89 | 499 |  |

8073
(1900) $22^{\mathrm{h} .} 25^{\mathrm{m}} 27^{\mathrm{s}}\left(+2^{\mathrm{s}} .22\right) ; \quad+57^{\circ} 54^{\prime} .2\left(+0^{\prime} .31\right)$

Period: $5^{\mathrm{d}} 8^{\mathrm{h}} 47^{\mathrm{m}} 39^{\mathrm{s}} .3-$; Variation: $3^{\mathrm{m}} .7-4^{\mathrm{m}} .9$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
| $\boldsymbol{\zeta}$ | 36 | $+57^{\circ} 2475$ | 0.0 | 3.7 |
| $\vdots$ | 40 | $65^{\circ} 1814$ | 0.8 | 3.7 |
| $\varepsilon$ | 38 | $56^{\circ} 2741$ | 4.6 | 4.2 |
| $\xi$ | 35 | $+63^{\circ} 1802$ | 5.7 | 4.4 |

Notes:
Owing to its short period, this star would require more systematic observations. The accurate knowledge of its period, however, will make these observations useful.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:




METHOD BY STEPS:

| Jan. | 28 | 12.2 | II | $\zeta 3 \delta 6 \varepsilon$ | 0.8 | 1.5 | 1.2 | 300 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Febr. | 12 | 14.8 | I | $\zeta 7 \delta 0 \varepsilon$ | 5.8 |  | 5.8 | 315 |  |
|  | 16 | 15.2 | II | $\zeta 5 \delta 18$ | 4.3 | 3.8 | 4.1 | 319 |  |
| Mar. | 13 | 14.8 | II | $\zeta 3 \delta 28$ | 2.8 | 2.8 | 2.8 | 344 | F G. |
|  | 17 | 14.5 | I | $\zeta 3 \mathrm{f} 18$ | 3.3 | 3.4 | 3.4 | 348 |  |
|  | 24 | 15.9 | II | $\zeta 2$ ¢ 3 \& | 1.8 | 1.8 | 1.8 | 355 | " |
|  | 28 | 15.3 | II | $\zeta 2 \delta 38$ | 1.8 | 1.8 | 1.8 | 359 | " |
| Apr. | 18 | 15.6 | II | $\zeta 2$ ¢ 2 ع | 2.3 | 2.3 | 2.3 | 380 | " |
|  | 20 | 15.3 | I | $\zeta 3 \delta 2 \varepsilon$ | 2.8 | 2.8 | 2.8 | 382 | " |
|  | 25 | 16.0 | I | ${ }_{\zeta} 3$ \% 2 ع | 2.8 | 2.8 | 2.8 | 387 | " |
| May | 15 | 15.8 | II | ${ }_{\zeta}^{6} 1{ }^{\circ} 1 \varepsilon$ | 2.3 | 2.3 | 2.3 | 407 | " |
|  | 19 | 15.9 | II | $\zeta 2 \mathrm{f} 2 \mathrm{\varepsilon}$ | 2.3 | 2.3 | 2.3 | 411 |  |
| June | 14 | 15.8 | II | $\zeta 2 \delta 2 \varepsilon$ | 2.3 | 2.3 | 2.3 | 437 |  |
|  | 17 | 16.5 | I | $\zeta 0.5{ }^{\circ} 38$ | 1.0 | 0.7 | 0.9 | 440 |  |
|  | 23 | 15.5 | II | ${ }_{5} 1$ ¢ 3 e | 1.3 | 1.1 | 1.2 | 446 |  |
| July | 6 | 16.6 | D | $\zeta 3$ \% 1 ¢ | 3.3 | 3.4 | 3.4 | 459 |  |
|  | 10 | 16.1 | I | ${ }_{5}^{5} 2{ }_{\text {¢ }} 3$ ¢ | 1.8 | 1.8 | 1.8 | 463 |  |
| Aug. | 15 | 16.8 | II | ${ }_{\zeta} 0{ }_{\delta} 4^{\varepsilon}$ | 0.3 |  | 0.3 | 499 |  |

*) The original had $\varepsilon 2 \zeta, \varepsilon 4^{\prime}$, corrected into $\zeta 2 \varepsilon, 44 \varepsilon$. The reading given above is more probable.

Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| $\rho$ | $\cdots \cdots$. | $+31^{\circ} 2628$ | 0.0 | 3.6 |
|  | $\cdots P$. |  |  |  |
| c | $\cdots \cdots$ | $+30^{\circ} 2536$ | 2.0 | 4.5 |
|  | $+27^{\circ} 2388$ | 6.1 | 5.9 |  |
|  |  |  |  |  |

## Notes:

This star was considered variable by Schmidt in 1867, but was not entered in Schönfeld's Catalogue II (1875). It is in all three of Chandler's catalogues under the designation $W$ Bootis, with the number 5274 . In the Potsdam Photometric Durchmusterung II (Bd. 13, 1899, Note to No. 2517), however, the star is pronounced: not variable. The following observations may serve to confirm this judgment. The letter $c$ for the third comparison star is not Bayer's.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:

| 84 Sept. | 9 | $55^{\text {h }} 5$ | I | $\rho 7 \mathrm{c}$ | 4.3 | 4.0 | 09429 | F G. used throughout. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 13 | 14.3 | III | $\sigma 4 \mathrm{c}$ <br> $\rho 88 \mathrm{c}$ | 3.6 4.9 | 4.5 | 433 |  |
|  |  |  |  | $\sigma 5 \mathrm{c}$ | 4.1 |  |  |  |
|  | 17 | 14.1 | I | $\rho 6 \mathrm{c}$ | 3.7 | 3.5 | 437 |  |
|  | 20 | 14.5 | I | $\circ$ $\mathrm{c}_{\text {c }}^{\text {c }}$ | 3.2 4.3 | 3.8 | 440 |  |
|  |  |  |  | $\sigma 3 \mathrm{c}$ | 3.2 |  |  |  |
| 85 May | 3 | 15.2 | II | $\rho 7 \mathrm{c}$ | 4.3 | 3.8 | 665 |  |
|  |  |  |  | $\sigma 3 \mathrm{c}$ | 3.2 |  |  |  |
|  | 8 | 15.2 | I | $\rho 6 \mathrm{c}$ | 3.7 | 3.7 | 670 |  |
| June | 18 | 15.6 |  | $\checkmark 4 \mathrm{c}$ | 3.6 4.3 | 3.6 | 711 |  |
|  |  | 15.6 | D | $\rho 7 \mathrm{c}$ <br> $\sigma$ | 4.3 2.8 | 3.6 | 71 |  |
| July | 2 | 15.6 | I | $\rho 9 \mathrm{c}$ ? | (5.5) | (4.8) | 725 | N. E. obs. doubtful. |
|  | 6 | 15.9 | III | $\begin{array}{r}\sigma \\ \rho 9 \mathrm{c} \\ \hline\end{array}$ | -5.5 | 4.6 | 729 |  |
| Sept. |  |  |  | $\sigma 4 \mathrm{c}$ | 3.6 |  |  |  |
|  | 14 | 15.4 | II D | $\rho 8 \mathrm{c}$ | 4.9 | 4.3 | 09799 |  |
|  |  |  |  | $\sigma 4 \mathrm{c}$ |  |  |  |  |
| 86 Apr. | 4 | 15.5 | II | $\rho 8 \mathrm{c}$ | 4.9 | 4.5 | 10001 |  |
|  | 6 |  |  | $\sigma 5$ c | 4.1 |  |  |  |
|  |  | 15.8 | I | $\rho 6 \mathrm{c}$ | 3.7 | 3.6 | 003 |  |
|  | 27 | 14.8 | II | $\sigma 3.5 \mathrm{c}$ $\rho 7 \mathrm{c}$ | 3.4 4.3 | 4.2 | 024 |  |
|  |  |  |  | ${ }_{\sigma} 5 \mathrm{c}$ | 4.1 |  |  |  |
| May | 1 | 15.5 | II | $\rho 7 \mathrm{c}$ | 4.3 | 3.8 | 028 |  |
|  |  |  |  | ¢ 3 c | 3.2 |  |  |  |
|  | 6 | 14.9 | III | $\rho 8 \mathrm{c}$ | 4.9 | 4.3 | 033 |  |
|  | 11 | 15.5 |  | $\sigma 4 \mathrm{c}$ | 3.6 | 2.8 | 038 |  |
|  |  | 15.5 | I 2 | $\rho$ $\sigma$ $\sigma$ 1 c | 2.4 | 2.8 | 038 |  |
|  | 18 | 15.5 | DDD | ¢ 63 | 3.7 | 3.3 | 045 |  |
|  | 27 | 15.3 | I | $\begin{aligned} & \sigma \\ & \sigma 2 \mathrm{c} \\ & \rho\end{aligned}$ | 2.8 4.9 | 4.3 | 054 |  |
|  |  |  |  | $\sigma 4 \mathrm{c}$ | 3.6 |  |  |  |
| June | 2 | 15.6 | I | $\rho 8 \mathrm{c}$ | 4.9 | 4.5 | 060 |  |
|  |  |  |  |  | 4.1 |  |  |  |
|  | 17 | 15.3 | I DD | $\rho 7 \mathrm{c}$ $\sigma 3 \mathrm{c}$ | 4.3 3.2 | 3.8 | 075 |  |

METHOD BY STEPS:

| 87 Mar. | 14.9 | I | $\rho 6 \mathrm{~W} 2 \mathrm{c}$ | 5.0 | 4.6 | 4.7 | 348 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 14.8 |  | $\sigma 3 \mathrm{~W} 2 \mathrm{c}$ | 4.5 | 4.5 |  |  |
|  |  | II | $\rho 4 W 4 \mathrm{c}$ | 3.0 | 3.0 | 3.1 | 352 |
|  |  |  | $\sigma 2 \mathrm{~W} 4 \mathrm{c}$ | 3.0 | 3.4 |  |  |



## 2 Serpentis

(1900) $18^{\mathrm{h}} 51^{\mathrm{m}} 15^{\mathrm{s}} \quad\left(+2^{\mathrm{s} .98)} ;+4^{\circ} 4^{\prime} .0 \quad\left(+0^{\prime} .08\right)\right.$

Relative brightuess $0.4-1.4$ ?
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| $\gamma$ | 63 | $+10^{\circ} 4043$ |  | 0.0 |
| $\beta$ | 70 | $+6^{\circ} 4357$ | 2.0 | 2.8 |
| $\mu$ | 52 | $+7^{\circ} 4132$ | 4.2 | 4.8 |

## Notes:

The numbers of the column ASV. refer to the chart of $\eta$ Aquilx (Ser. V, Charta XIV), as the comparison stars of $\vartheta$ Serpentis and $\eta$ Aquilæ are the same. The star is double, and the relative brightness of the two components is suspected of variability (See Potsdam Photom. Durchmusterung I, p. 482, Note to 2610-11). In the following observations the two components were estimated as one star.

| $1800+$ | Gr. M. T. Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:


## METHOD BY STEPS:

| 87 May | 18 | 16.7 | I |  | 2.6 2.0 | 2.5 | 2.4 | 410 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| June | 14 | 15.8 | II | r1\% 2.5 \% | 1.3 | 1.2 | 1.5 | 437 |
|  | 16 | 15.8 | II |  | 2.0 2.1 | 2.1 | 2.7 | 439 |
|  |  |  |  | ¢ 2 \% | 4.0 |  |  |  |
|  | 17 | 16.4 | I | ¢ 3 \% 2 " | 2.6 | 2.5 | 2.7 | 440 |
|  | 23 | 15.7 | II |  | 3.0 2.1 | 2.1 | 2.1 | 46 |
|  |  |  |  | ¢0\%* | 2.0 |  | 2.1 | 46 |
| July | 10 | 16.1 | I | r3*2\% | 2.6 | 2.5 | 2.7 | 463 |
|  |  |  |  | $\beta 1$ \% | 3.0 |  |  |  |
|  | 18 | 16.5 | II | \% 3 \% $5 \mu$ | 1.1 | 1.6 | 1.1 | 471 |
| Aug. |  |  |  |  | 0.5 |  |  |  |
|  | 15 | 16.7 | II |  | 3.1 3.0 | 2.8 | 3.0 | 499 |

(1900) $20^{\mathrm{h}} 14^{\mathrm{m}} 6^{\mathrm{s}}\left(+2^{\mathrm{s}} .21\right) ;+37^{\circ} 43^{\prime} .3\left(+0^{\prime} .18\right)$

Magnitude: PD $=5^{\mathrm{x}} .0, \mathrm{HP} .=4^{\mathrm{x}} .9$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| $\mathrm{b}^{2}$ | 31 | $+36^{\circ} 3907$ | 0.0 | 4.8 |
| $\mathrm{~b}^{3}$ | 34 | $+36^{\circ} 3955$ | 1.0 | 5.1 |
| c | 37 | $+36^{\circ} 3998$ | 3.0 | 5.5 |

## Notes:

Although no variations have been established in the brightness of this star for the last two hundred years, it is in all the catalogues of Schänfeld and Chandler, because it was subject to considerable fluctuations in the seventeenth century. The following observations may be of use to future discussions of its variability. The numbers in the column ASV. refer to Charta XV, Series V.

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |

DECIMAL METHOD:

| 84 Sept. | 11 | 15.5 | I | $b^{2} 5 c$ <br> $b^{2} 10 b^{3}$ | 1.5 | 1.3 | 09431 | F G. used throughout. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 13 | 14.7 | III | $\mathrm{b}^{2} 7 \mathrm{c}$ | 2.1 | 2.1 | 433 |  |
|  |  |  |  | $\mathrm{b}^{3} 5 \mathrm{c}$ | 2.0 |  |  |  |
|  | 17 | 15.6 | I | $\mathrm{b}^{2} 5 \mathrm{c}$ | 1.5 | 2.0 | 437 |  |
|  |  |  |  | $\mathrm{b}^{3} 7 \mathrm{c}$ | 2.4 |  |  |  |
|  | 24 | 15.7 | I | $\mathrm{b}^{2} 7 \mathrm{c}$ $\mathrm{b}^{3} 9 \mathrm{c}$ | 2.1 1.8 | 2.0 | 444 |  |
| Oct. | 10 | 15.3 | I | $\mathrm{b}^{\circ} 7 \mathrm{c}$ | 2.1 | 1.5 | 460 |  |
|  |  |  |  | $\mathrm{b}^{2} 9 \mathrm{~b}^{3}$ | 0.9 |  |  |  |
|  | 12 | 15.6 | I | $\mathrm{b}^{2} 4 \mathrm{c}$ | 1.2 | 1.0 | 462 |  |
|  |  |  |  | $\mathrm{b}^{2} 8 \mathrm{~b}^{3}$ | 0.8 |  |  |  |
|  | 16 | 15.4 | II | $\mathrm{b}^{2} 5 \mathrm{c}$ | 1.5 | 1.9 | 466 |  |
|  |  |  |  | - ${ }^{3} 6 \mathrm{c}$ | 2.2 |  |  |  |
| Nov. | 7 | 15.4 | I | $\mathrm{b}^{2} 5 \mathrm{c}$ | 1.5 | 1.7 | 488 |  |
|  |  |  |  | $\mathrm{b}^{3} 4 \mathrm{c}$ | 1.8 |  |  |  |
|  | 18 | 15.0 | I | ${ }^{\mathrm{b}^{2} 6}{ }^{3} \mathrm{c}$ | 1.8 | 1.9 | 499 |  |
| Dec. | 9 | 14.7 | II | $\mathrm{b}^{2} 7 \mathrm{c}$ | 2.1 | 2.2 | 520 |  |
|  |  |  |  | $\mathrm{b}^{3} 6 \mathrm{c}$ | 2.2 |  |  |  |
|  | 18 | 15.0 - | I | $\mathrm{b}^{2} 9 \mathrm{c}$ | 2.7 | 2.7 | 529 |  |
|  |  |  |  | $\mathrm{b}^{3} 8 \mathrm{c}$ | 2.6 |  |  |  |
| 85 Jan. | 9 | 14.4 | I | $\mathrm{b}^{2} 8$ c | 2.4 | 2.4 | 551 |  |
|  |  |  |  | $\mathrm{b}^{3} 7 \mathrm{c}$ | 2.4 |  |  |  |
| June | 4 | 15.9 | 1 | $\mathrm{b}^{2} 2 \mathrm{c}$ | 0.6 | 1.1 | 697 |  |
|  |  |  |  | $\mathrm{b}^{3} 3 \mathrm{c}$ | 1.6 |  |  |  |
|  | 8 | 16.0 | I | $\mathrm{b}^{2} 0 \mathrm{c}$ | 0.0 | 0.6 | 701 |  |
|  | 15 | 15.7 | III | $\mathrm{b}^{3} 1$ 1 $\mathrm{~b}^{2}$ 0 c | 1.2 0.0 | 0.6 | 708 |  |
|  | 15 | 15.7 | H | $\mathrm{b}^{3} 1 \mathrm{c}$ | 1.2 |  |  |  |
|  | 17 | 15.6 | D | $\mathrm{b}^{2} 1 \mathrm{c}$ | 0.3 | 0.9 | 710 |  |
|  |  |  |  | $\mathrm{b}^{3} 2 \mathrm{c}$ | 1.4 |  |  |  |
| July | 6 | 16.1 | III | $\mathrm{b}^{2} 2 \mathrm{c}$ ? | 0.6 | 1.1 | 729 |  |
|  |  |  |  | $\mathrm{b}^{3} 3 \mathrm{c}$ ? | 1.6 |  |  |  |
| Sept. | 14 | 15.6 | 11) | b $\mathrm{b}^{3} 5 \mathrm{c}$ 4 c | 1.5 | 1.7 | 799 |  |
| Oct. | 1 | 15.2 | II | $\mathrm{b}^{2} 5 \mathrm{c}$ | 1.5 | 1.6 | 816 |  |
|  |  |  |  | $\mathrm{b}^{3} 3 \mathrm{c}$ | 1.6 |  |  |  |
|  | 5 | 15.3 | I | $\mathrm{b}^{2} 4 \mathrm{c}$ | 1.2 | 1.4 | 820 |  |
|  |  |  |  | $\mathrm{b}^{3} 3 \mathrm{c}$ | 1.6 |  |  |  |
|  | 9 | 15.6 | I | $\mathrm{b}^{2} 4 \mathrm{c}$ | 1.2 | 1.3 | 824 |  |
|  |  |  |  | $\mathrm{b}^{3} 2 \mathrm{c}$ | 1.4 |  |  |  |
| Nov. | 9 | 15.0 | I | $\mathrm{b}^{2} 4 \mathrm{c}$ | 1.2 | 1.3 | 855 |  |
| Dec. | 6 | 15.0 | I | $b^{3} 2 \mathrm{c}$ $\mathrm{b}^{2} 6 \mathrm{c}$ | 1.4 | 1.8 | 882 |  |
|  |  |  |  | $\mathrm{b}^{3} 4 \mathrm{c}$ | 1.8 |  |  |  |
|  | 11 | 15.2 | I | $\mathrm{b}^{2} 5 \mathrm{c}$ | 1.5 | 1.7 | 09887 |  |
|  |  |  |  | $\mathrm{b}^{3} 4 \mathrm{c}$ | 1.8 |  |  |  |


| $1800+$ | Gr. M. T. | Sky | Comparisous | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 86 May | 15.9 | I D | $\mathrm{b}^{2} 2 \mathrm{c}$ |  | 0:6 | 0.9 | 10045 |  |
|  | 15.8 | I D | $b^{3} 1 \mathrm{c}$ $\mathrm{b}^{2} 3 \mathrm{c}$ |  | 1.2 | 1.2 | 054 |  |
|  |  |  | $b^{3} 2 \mathrm{c}$ |  | 1.4 |  |  |  |
|  | 15.8 | III | $\mathrm{b}^{2} 4 \mathrm{c}$ |  | 1.2 | 1.6 | 056 |  |
|  |  |  | $\mathrm{b}^{3} 5 \mathrm{c}$ |  | 2.0 |  |  |  |
| June | 15.4 | I | $\mathrm{b}^{2} 3 \mathrm{c}$ |  | 0.9 | 1.4 | 061 |  |
|  |  |  | $\mathrm{b}^{3} 4 \mathrm{c}$ |  | 1.8 |  |  |  |
|  | 15.4 | ID | $b^{3} 1 \mathrm{c}$ $\mathrm{b}^{3} 2 \mathrm{c}$ |  | 1.3 1.4 | 0.9 | 075 |  |
| Oct. | 14.6 | I D | $\mathrm{b}^{2} 3 \mathrm{c}$ |  | 0.9 | 1.4 | 181 |  |
|  | 14.5 | II | b ${ }^{2} 1 \mathrm{c}$ |  | 1.8 | 0.9 | 198 |  |
|  |  |  | $\mathrm{b}^{3} 2 \mathrm{c}$ |  | 1.4 |  |  |  |
|  | 15.0 | I | $\mathrm{b}^{2} 2 \mathrm{c}$ |  | 0.6 | 1.1 | 209 |  |
| Nov. 25 | 13.5 | II | b $i^{2}$ 2 c |  | 1.6 0.6 | 1.1 | 236 |  |
| Dec. $\begin{aligned} & 15 \\ & \\ & 29\end{aligned}$ |  |  | $\mathrm{b}^{3} 3 \mathrm{c}$ |  | 1.6 |  |  |  |
|  | 14.1 | I | $\mathrm{b}^{3} 4 \mathrm{c}$ |  | 1.8 | 1.4 | 256 |  |
|  | 14.2 | I | $\mathrm{b}^{3} 5 \mathrm{c}$ |  | 2.0 | 1.6 | 270 |  |
|  |  |  | $\mathrm{b}^{2} 4 \mathrm{c}$ |  | 1.2 |  |  |  |

METHOD BY STEPS:


## III. Observations made by J. G. Hagen. S. J., from 1888 to 1890.

The following observations were made at the Georgetown College Observatory, with a 5 -inch equatorial by Troughton \& Simms. The eye-piece had a power of 50 diameters, with a field of less than a degree, rather too small for the purpose. The observing list consisted almost exclusively of southern variables, which could not be well observed in more northern latitudes. Most of the results of these observations were published at the time in the Astronomical Journal.

806 o Ceti Series IV \& V.
(1900) $2^{\mathrm{n}} 14^{\mathrm{m}} 18^{\mathrm{s}}\left(+3^{\mathrm{s}} .03\right) ;-3^{\circ} 25^{\prime} .7 \quad\left(+0^{\prime} .27\right)$

Period: $331^{\mathrm{d}} .6$; Variation: $2^{\mathrm{M}}-9^{\mathrm{n}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| 1 |  | $-3^{\circ} 340$ | 0.0 | [7.7] BD. |
| m | 2 | -4*379 | 6.0 | 8.0 |
| n | 5 | $-2^{\circ} 396$ | 13.7 | 8.5 |
| p | 6 | $-3^{\circ} 363$ | 15.8 | 8.6 |
| r | 7 | $-3^{\circ} 355$ | 22.8 | 8.8 |
| q | ...... | $-3^{\circ} 362$ | 24.8 | [9.2] BD. |

## Notes:

In order to avoid confusion in the notation, the variable was designated by M , instead of the Greek letter $\%_{0}$ These few observations were made incidentally in preparing the two charts for the Atlas. The numbers under ASV. and Magn. (except those in parenthesis) are taken from the IV. Series of the Atlas.


814 S Persei Series III.
(1900) $2^{\mathrm{h}} 15^{\mathrm{m}} 41^{s} .\left(+4^{4} .27\right) ;+58^{\circ} 7^{\prime} .8 \quad\left(+0^{\prime} .28\right)$

Variation: $8 \frac{1}{2}^{\mathrm{M}}-12^{\mathrm{M}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| ---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
| g | 1 | $+58^{\circ} 471$ | 0.0 |  |
| f | 4 | $+55^{\circ} 467$ | 3.3 | 8.8 |
| e | 5 | $+58^{\circ} 452$ | 4.3 | 8.2 |
| k | 7 | $+57^{\circ} 549$ | 5.3 | 8.5 |
| b | 18 | $+57^{\circ} 557$ | 14.5 | 9.5 |
| c | 36 | $\ldots \cdots \cdots$ | 19.1 | 10.9 |

Notes:
These observations are a continuation of those made from 1883 to 1888 , and were reduced on the same scale as the latter, although the instrument was a different one.


845
R Ceti
Series I.
(1900) $2^{\mathrm{n}} 20^{\mathrm{m}} 55^{\mathrm{c}}\left(+3^{s} .06\right) ;-0^{\circ} 37^{\prime} .8 \quad\left(+0^{\prime} .27\right)$

Period: 167 $7^{\mathrm{d}} 0$; Variation: $8^{\mathrm{M}}-13 \frac{1}{2}^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| e | 1 | $-1^{\circ} 338$ | 0.0 | 8.0 |
| a | 2 | $-0^{\circ} 267$ | 10.2 | 8.5 |
| d | 3 | $-0^{\circ} 365$ | 14.2 | 8.6 |
| f | 4 | $-0^{\circ} 363$ | 23.7 | 9.1 |
| c | 6 | $-1^{\circ} 333$ | 27.1 | 9.4 |
| b | 7 | $-1^{\circ} 339$ | 28.1 | 9.4 |
| L | 9 | $\ldots \ldots \ldots$. | 34.6 | 9.8 |


(1900) $2^{\mathrm{h}} 28^{\mathrm{m}} 56^{\mathrm{s}}\left(+2^{\mathrm{s}} .88\right) ;-13^{\circ} 35^{\prime} .2\left(+0^{\prime} .27\right)$

Period: $235^{\mathrm{d}} .8$; Variation: $7^{\mathrm{N}}-12^{\mathrm{N}}$.
Comparison Stars :

| Obs. | ASV. | BD . | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| A | 1 | $-12^{\circ} 481$ | 0.0 | 6.8 |
| i | 2 | $13^{\circ} 492$ | 6.7 | 7.5 |
| n | - | $13^{\circ} 493$ | 9.7 | (8.8) BD. |
| m | 3 | $12^{\circ} 469$ | 10.8 | 8.0 |
| h | - | $12^{\circ} 489$ | 14.2 | (9.0) |
| , | - | $13^{\circ} 473$ | 18.8 | (8.9) |
| g | 8 | $13^{\circ} 483$ | 22.4 | 8.7 |
| f | 7 | $13^{\circ} 481$ | 24.9 | 8.6 |
| B | 11 | $13^{\circ} 487$ | 34.9 | 9.2 |
| e | 10 | $14^{\circ} 479$ | 35.2 | 9.0 |
| d | 12 | $13^{\circ} 474$ | 41.6 | 9.3 |
| a | 17 | $13^{\circ} 476$ | 46.6 | 9.6 |
| b | 19 | $-13^{\circ} 478$ | 50.6 | 9.8 |
| c | 22 | ......... | (53.6) | 10.0 |

\begin{tabular}{|c|c|c|c|c|c|c|c|c|c|}
\hline 1800 \& \& Gr. M. T. \& Sky \& Comparisons \& I \& .II \& Mean \& $2400000+$ \& Remarks <br>
\hline \multirow[t]{7}{*}{89 Dec.} \& 12 \& 13.5 \& I \& 1) 1 U 2 c \& 51.6 \& 51.6 \& 51.6 \& 11349 \& <br>
\hline \& 20 \& 15 \& III \& d 3 U 5 b 1 a \& 45.1
45.6 \& 45.3 \& 45.3 \& 357 \& <br>
\hline \& 21 \& 14 \& I \& d2 U 3 a \& 43.6 \& 43.7 \& 44.5 \& 358 \& <br>
\hline \& \& \& \& d 2 U 4 c \& (46.6) \& (45.6) \& \& \& <br>
\hline \& 22 \& 14 \& I \& d 2 U 4 a \& 43.1 \& 43.3 \& 43.2 \& 359 \& <br>
\hline \& 24 \& 13.5 \& III \& d2 U3a2c \& 43.6 \& 43.7 \& 43.7 \& 361 \& <br>
\hline \& 26 \& 12.5 \& I \& e 4 U 2 d \& 38.4 \& 39.5 \& 39.0 \& 363 \& <br>
\hline \multirow[t]{11}{*}{90 Jan.} \& 12 \& 13.3 \& III \& f 3 U 3 e 4 d \& 30.0 \& 30.2 \& 30.1 \& 380 \& <br>
\hline \& 13 \& 15.5 \& I \& f 3 U 4 e \& 29.5 \& 29.3 \& 29.4 \& 381 \& <br>
\hline \& 16 \& 15 \& I \& in 3 U 3 g
h 3 U 5 f \& 18.3
18.6 \& 18.3
18.2 \& 18.4 \& 384 \& <br>
\hline \& 18 \& 13.3 \& III \& h2 U5f \& 18.0 \& 17.3 \& 16.8 \& 386 \& <br>
\hline \& \& \& \& i5U3g2f,i5h \& 15.5 \& 16.5 \& \& \& <br>
\hline \& 21 \& 15.5 \& I \& A 4 U1i \& 4.8 \& 5.4 \& 5.1 \& 389 \& <br>
\hline \& 22 \& 15 \& I \& A 4 U0i \& 5.3 \& \& 5.3 \& 390 \& <br>
\hline \& 23 \& 14.5 \& I \& i 2 U 3 m \& 8.2 \& 8.3 \& 8.3 \& 391 \& <br>
\hline \& 24 \& 13.5 \& III D \& i 3 U 0 m \& 10.2 \& \& 10.2 \& 392 \& Near horizon. <br>
\hline \& 27 \& 13.5 \& I D \& A 5 U ${ }^{\text {i }}$ ( 5 m \& 5.0
6.8 \& 7.4 \& 6.4 \& 395 \& <br>
\hline \& 28. \& 12.8 \& DDD \& A 5 U 1 i \& 5.3 \& 5.6 \& 5.2 \& 396 \& <br>
\hline \multirow[t]{5}{*}{Feb.} \& \& 13 \& I \& A 5 U 6 mm \& 4.9
4.9 \& 4.9
5.4 \& 5.2 \& 409 \& <br>
\hline \& 12 \& 13 \& \& A 5 U 1 i \& 5.4 \& 5.6 \& 5.5 \& 411 \& <br>
\hline \& 16 \& $13^{\text {h }}-15^{\text {n }}$ \& III \& i 2 U 8 g ! \& 11.5 \& 9.8 \& 10.7 \& 415 \& <br>
\hline \& 18 \& / \& 111 \& i 2 U 8 g \& 11.5 \& 9.8 \& 10.7 \& 417 \& Difficult. <br>
\hline \& 22 \& " 8 \& III D \& i 3 U 6 g \& 13.0 \& 11.9 \& 12.5 \& 421 \& Near horizon. <br>
\hline \multirow[t]{2}{*}{Mar.} \& $$
\begin{aligned}
& 6 \\
& 6 \\
& 8
\end{aligned}
$$ \& $$
\underset{13^{\mathrm{h}}-12.8}{5^{\mathrm{n}}}
$$ \& I DDD \& i 3 U 7 g
i 10 U 2 g \& 12.5 \& 11.4
19.7 \& 12.0
19.1 \& 433
435 \& Near horizon. <br>
\hline \& \& \& \& $110 \mathrm{U}^{2 \mathrm{~g}}$ \& \& \& \& \& <br>
\hline \multirow[t]{3}{*}{Aug. Sept.} \& 25 \& 17.5 \& II D \& g 1 U 3 f \& 22.6 \& 23.0 \& 22.8 \& 605 \& <br>
\hline \& +888 \& 17.1
16.3 \& ${ }_{\text {III }}^{\text {II }}$ \& i 7 U 41
i 3 U 101 \& 13.8
10.5 \& 13.7
9.7 \& 13.8
10.1 \& 619
626 \& <br>
\hline \& 17 \& 16.3
15.3 \& I \& 13 4 U $\mathrm{U}^{101}$ \& 10.5 \& 9.7 \& 10.7 \& 626
628 \& <br>
\hline \multirow[t]{7}{*}{Oct.

Nov.} \& 3 \& 14 \& I \& $\mathrm{U}=\mathrm{i}$ \& 6.7 \& \& 6.7 \& 644 \& <br>
\hline \& 5 \& 13 \& III \& i 1 U ! \& 7.7 \& \& 7.7 \& 646 \& <br>
\hline \& 8 \& 14.3 \& I \& $\mathrm{U}=\mathrm{i}$ ! \& 6.7 \& \& 6.7 \& 649 \& <br>
\hline \& 17 \& 13.5 \& I \& A 7 U 2 i \& 5.8 \& 5.2 \& 5.5 \& 658 \& <br>
\hline \& 21 \& 16.5 \& I \& A 6 U 2 i \& 5.3 \& 5.0 \& 5.2 \& 662 \& <br>
\hline \& 31 \& 12.5 \& III \& i 3 U 0 n ! \& 9.7 \& \& 9.7 \& 672 \& <br>
\hline \& 3 \& 15 \& III \& $\mathrm{i}_{5} 5 \mathrm{U} 12 \mathrm{~g}$ \& 11.0 \& 11.3 \& 10.8 \& 675 \& <br>
\hline \multirow{9}{*}{Nov.} \& 7 \& $12^{\text {n }}-15^{\text {b }}$ \& II \& $\mathrm{U}_{\mathrm{i} 6 \mathrm{U}} \mathbf{4} 8 \mathrm{~g}$ \& 10.2
12.7 \& \& 11.9 \& 679 \& <br>
\hline \& \& 12-15 \& \& n 3 U 4 h \& 11.5 \& 11.6 \& \& \& <br>
\hline \& 13 \& $13^{\mathrm{h}}-14^{\mathrm{b}}$ \& II \& i 10 U 2 h 10 g \& 14.5 \& 12.9 \& 1.3.8 \& 685 \& <br>
\hline \& \& \& \& ${ }^{1} 7 \mathrm{U} 2 \mathrm{~h}$ \& 14.5 \& 13.2 \& \& \& <br>
\hline \& 18. \& 15.3 \& I D \& n10 10 U 8 g
h 2 U \& 17.0 \& 16.8 \& 16.4 \& 690 \& <br>
\hline \& \& \& \& h 1.5 U \& 15.7 \& \& \& \& <br>
\hline \& \& \& I DDD \& g 2f 4 U5 B \& 29.4 \& 29.3 \& 29.4 \& 700 \& <br>
\hline \& 30 \& $13^{\text {b }}-15^{\text {n }}$ \& \& f7 U4B1e \& 31.4 \& 31.3 \& 31.4 \& 702 \& <br>

\hline \& \& \& \& $$
\begin{aligned}
& f 1 \mathrm{~g} 3 \mathrm{l} \\
& \mathrm{i} 7 \mathrm{~m} 7 \mathrm{~g}, \mathrm{~h} 3 \mathrm{~m}
\end{aligned}
$$ \& \& \& \& \& <br>

\hline
\end{tabular}

| $1800+$ | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{array}{lll}90 & \text { Dec. } & 10 \\ 13\end{array}$ | $\begin{gathered} 13^{\mathrm{h}}-15^{\mathrm{h}} \\ 14 \end{gathered}$ | $\frac{1}{\text { I }}$ | $\begin{aligned} & \mathrm{e} 4 \mathrm{U} 3 \mathrm{~d} \\ & \mathrm{U}=\mathrm{d}! \end{aligned}$ | 38.9 41.6 | 38.9 | $\begin{aligned} & 38.9 \\ & 41.6 \end{aligned}$ | $\begin{array}{r} 11712 \\ 715 \end{array}$ | 3-in glass, difficult. |

(1900) $12^{\mathrm{h}} 14^{\mathrm{m}} 27^{\mathrm{s}}\left(+3^{\mathrm{s}} .10\right) ;-18^{\circ} 42^{\prime} .0 \quad\left(-0^{\prime} .33\right)$

Period: $318^{\mathrm{d}} .5$; Variation: $7^{\mathrm{N}}-12^{\mathrm{N}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| a | 1 | $-18^{\circ} 3379$ | 2.0 | 7.4 |
| p | - | $18^{\circ} 3380$ | 5.3 | (7.0) BD. |
| b | 2 | $18^{\circ} 3368$ | 14.6 | 7.8 |
| c | 3 | $18^{\circ} 3369$ | 19.3 | 8.0 |
| r | - | $19^{\circ} 3466$ | 24.3 | (8.5) BD. |
|  | - | $17^{\circ} 3596$ | 27.0 | (8.3) BD. |
| d | 8 | $18^{\circ} 3372$ | (27) | 9.5 |
| m | 4 | $18^{\circ} 3362$ | 31.0 | 8.5 |
| k | 5 | $18^{\circ} 3365$ | 31.3 | 8.8 |
| n | 6 | $18^{\circ} 3373$ | 31.3 | 8.9 |
| h | 7 | $18^{\circ} 3364$ | 34.3 | 9.2 |
| e | 11 | $-18^{\circ} 3366$ | (35) | 10.0 |


*) The original has: p 2 R 1 d, which seems to be an error in recording.
(1900) $12^{\mathrm{h}} 20^{\mathrm{m}} 52^{\mathrm{s}} \quad\left(+3^{\mathrm{s}} .09\right) ;-2^{\circ} 51^{\prime} .5 \quad\left(-0^{\prime} .31\right)$

Period: $17^{\mathrm{d}} .2711$; Variation: $9^{\mathrm{N}}-10^{\mathrm{n}}$.
Comparison Stars :

| Obs. | ASV. | BD. |  | Steps |
| ---: | ---: | ---: | :---: | :---: |
|  |  |  | Magn. |  |
| f | 5 | $-3^{\circ} 3458$ | 6.0 | 8.3 |
| g | - | $3^{\circ} 3461$ | 10.0 | $(8.5) \mathrm{BD}$. |
| h | 8 | $3^{\circ} 3460$ | 16.9 | 9.0 |
| c | - | $2^{\circ} 3697$ | $(19.0)$ | $(8.9)$ |
| m | 24 | $2^{\circ} 3679$ | 21.4 | 9.5 |
| e | 18 | $2^{\circ} 3678$ | 23.6 | 9.4 |
| n | 23 | $3^{\circ} 3463$ | 23.9 | 9.6 |
| d | 15 | $2^{\circ} 3688$ | 27.0 | 9.5 |
| p | - | $1^{\circ} 2821$ | 29.7 | $(9.5)$ |
| b | 24 | $-2^{\circ} 3687$ | 31.3 | 9.7 |


*) The original has $f 2 \mathrm{~W} 4 \mathrm{e}$, which is a poor observation, the interval $e-f$ being too large. $c 2 \mathrm{~W}$ is taken from the second observation.
${ }^{* *}$ ) The original has $f 5 \mathrm{~W}$ instead of $c 2 \mathrm{~W}$; see preceding note.


4816 V Virginis Series 1.
(1900) $13^{\mathrm{h}} 22^{\mathrm{m}} 38^{\circ}\left(+3^{s} .09\right) ;-2^{\circ} 39^{\prime} .2\left(-0^{\prime} .31\right)$

Period: $250^{\text {a }} .5$; Variation: $8 \frac{1}{2}^{\mathrm{N}}-<13^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| i | 2 | $-2^{\circ} 3689$ | 15.7 | 9.1 |
| h | - | $3^{\circ} 3460$ | 16.9 | (8.9) BD. |
| k | 3 | $2^{\circ} 3690$ | 25.7 | 9.5 |
| d | 4 | $2^{\circ} 3688$ | 27.0 | 9.6 |
| b | 6 | $-2^{\circ} 3687$ | 31.3 | 9.9 |

## Notes:

These few observations were made incidentally with those of W Virginis, and were reduced by the same scale as the latter.

$(1900) 13^{\mathrm{b}} 27^{\mathrm{m}} 47^{\mathrm{s}}\left(+3^{\mathrm{s}} .13\right) ;-6^{\circ} 40^{\prime} .8 \quad\left(-0^{\prime} .31\right)$
Period: $376^{\mathrm{d}} .4$; Variation: $7^{\mathrm{M}}-12 \frac{1}{2}^{\mathrm{N}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| ---: | ---: | :---: | :---: | :---: |
|  |  |  |  |  |
| g | 1 | $-5^{\circ} 3706$ | 0.0 | $(6.6) \mathrm{BD}$. |
| c | 2 | $6^{\circ} 3839$ | 10.0 | 7.3 |
| f | 3 | $6^{\circ} 3843$ | 21.1 | 8.0 |
| d | 4 | $6^{\circ} 3834$ | 27.2 | 8.5 |
| e | 6 | $6^{\circ} 3840$ | 36.2 | 8.8 |
| h | 5 | $6^{\circ} 3832$ | 39.2 | 8.8 |
| k | 7 | $6^{\circ} 3836$ | 45.2 | 9.2 |
| n | 8 | $6^{\circ} 3833$ | 47.2 | 9.4 |
| m | 12 | $-6^{\circ} 3835$ | 53.2 | 9.8 |
|  |  |  |  |  |

## Notes:

The first three of these observations have a different scale of steps from those of the following year, the interval d-c being on the average only $S$ units, or less that one-half the same interval in the general scale. This will explain the discrepancies in their reduction to the latter scale.


6132 |  | R Ophinchi | Series I. |
| :---: | :---: | :---: |
| $(1900) 17^{\mathrm{n}} 2^{\mathrm{m}} 1^{\mathrm{s}}$ | $\left(+3^{\mathrm{s}} .44\right) ;-15^{\circ} 57^{\prime} .6$ | $\left(-0^{\prime} .08\right)$ |

Period: 302 $2^{4}$; Variation: $7 \frac{11}{2}^{n}-12^{N}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| a | $\ldots \ldots .$. | $-16^{\circ} 4434$ | 10.0 | $(7.5) \mathrm{BD}$. |
| b | $\ldots .$. | $16^{\circ} 4436$ | 16.1 | $(7.3)$ |
| c | 2 | $15^{\circ} 4466$ | 24.7 | 8.0 |
| d | 3 | $-16^{\circ} 4426$ | 33.7 | 8.2 |

## Notes:

When the variable is brighter than the brightest comparison star, as on September 9, 15, and 17, both forn ulas I and II can be used with a little modification. Thus on September 9 formula I requires 2 steps to be sul tracted from $a$ and 8 from $b$. Formula II only requires an interchange of $R$ and $a$, and will read thus:

$$
10=\mathrm{R}+2 \frac{16.1-\mathrm{R}}{8} \text {, or } \mathrm{R}=8.0
$$

and similarly for September 15 and 17.


6905 R Sagittarii Series I.
(1900) $19^{\mathrm{h}} 10^{\mathrm{m}} 49^{\mathrm{s}}\left(+3^{3} .52\right) ;-19^{\circ} 29^{\prime} .0 \quad\left(+0^{\prime} .10\right)$

Period: 268 $8^{\mathrm{d}} .7$; Variation: $7 \frac{1}{2}^{\mathrm{M}}-12 \frac{1}{2}^{\mathrm{M}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| g | 2 | $-19^{\circ} 5387$ | 0.0 | 8.0. |
| h | 3 | 5398 | 6.7 | 8.1 |
| i | 5 | 5388 | 15.5 | 8.4 |
| k | 8 | 5386 | 25.7 | 8.7 |
| 1 | 10 | 5375 | 28.7 | 8.8 |
| m | 12 | 5384 | 33.7 | 9.0 |
| n | 21 | $-19^{\circ} 5368$ | 42.2 | 9.5 |

## Notes:

The interval $l-k$ is not determined by the observations, and was taken from the ASV., the scale of which closely agrees with the steps of these observations.


6921
S Sagittarii

## Series I.

(1900) $19^{\mathrm{h}} 13^{\mathrm{m}} 35^{\mathrm{s}} \quad\left(+3^{\mathrm{s}} .51\right) ; \quad-19^{\circ} 12^{\prime} .4 \quad\left(+0^{\prime} .11\right)$

Period: $230^{\mathrm{d}} .6$; Variation: $10^{\mathrm{M}}-<14^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps. | Magn. |
| :--- | ---: | ---: | ---: | :---: |
| q | 30 | $\ldots \ldots \ldots$. | 0.0 | 10.0 |
| p | 35 | $-19^{\circ} 5397$ | 6.3 | 10.2 |

Notes:
This variable is too faint for the instrument used, and was observed only because it is in the same field as R Sagittarii, and happened to reach its maximum brightness at the same time with the latter.

*) The limit 15 may be concluded from the last three observations (November 5-13).

7468
T Aquarii
Series I.

Period: 203 ${ }^{4} .3$; Variation: $7 \frac{1}{2}^{\mathrm{N}}-13^{\mathrm{N}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps. | Magn. |
| ---: | ---: | ---: | ---: | :---: |
|  |  |  |  |  |
| b | 6 | $-5^{\circ} 5396$ | 0.0 | 7.9 |
| d | 7 | 5383 | 4.0 | 8.2 |
| c | 9 | 5393 | 9.7 | 8.5 |
| h | 10 | 5394 | 18.7 | 9.0 |
| k | - | 5385 | 20.0 | $(9.0) \mathrm{BD}$. |
| e | 22 | 5398 | 20.7 | 10.0 |
| f | 15 | 5387 | 25.7 | 9.3 |
| g | 21 | $-5^{\circ} 5389$ | 33.7 | 10.0 |


| 1800 |  | Gr. M. 'T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 89 Aug. | 31 | 15.5 | 1. | b 1 T 2 c | 4.3 | 3.2 | 3.8 | 11246 |  |
| Sept. | 19 | 15 |  | b 2 T 3 c | 4.3 | 3.9 | 4.1 | 265 |  |
|  | 21 | 9.2 | I | b 3 T 2 c | 5.3 | 5.6 | 5.5 | 267 |  |
|  | 27 | 8.2 | I | b 4 T 1 c | 6.3 | 7.8 | 7.0 | 273 |  |
| Oct. | 10 | 13.2 | III DDD | d 4 T T 3 e | 6.9 15.7 | 6.9 16.0 | 15.9 | 286 |  |
| 90 Sept. | 15 | 14.3 | I | f 3 T 5 g | 28.7 | 28.7 | 28.7 | 626 |  |
|  | 17 | 15 | 1 | f 2 T 6 g | 27.7 | 27.7 | 27.7 | 628 |  |
|  | 24 | 15 | I D2 | h 5 T 3 f | 23.2 | 23.1 | 23.2 | 635 |  |
|  | 26 | 12.8 | I DDD | h3 T6f | 20.7 | 21.1 | 20.9 | 637 |  |
| Oct. | 3 | 14 |  | d 4 T 1 c 5 h | 8.3 | 8.6 | 8.5 | 644 |  |
|  | 5 | 12.8 | III | d 5 c 2 T 10 h | 10.2 | 11.2 | 10.7 | 646 | $T<c!$ |
|  | 8 | 14 | I | d 5 c 2 T 8 h 3 f | 10.7 | 11.5 | 11.2 | 649 |  |
|  | 14 | 13.3 | III | d 2 T 4 c | 5.8 | 5.9 | 5.9 | 655 |  |
|  | 17 | 13.3 | I | d 5 T 3 c | 7.8 | 7.6 | 7.7 | 658 |  |
|  | 21 | 16 | I | d 3T6c | 5.3 | 5.9 | 5.6 | 662 | Near horizon. |
|  | 31 |  |  | d4c3T7! | 15.7 | 14.5 | 15.1 | 672 |  |
| Nov. | 3 5 | 14.8 13.1 | III |  | 7.8 9.3 | 7.6 8.9 | 7.7 9.1 | 675 677 |  |
|  | 9 | 13 | I | c 2 T 6 k 3 f | 12.9 | 12.3 | 12.6 | 681 |  |
|  | 13 | 13.1 | II | c 3.5 T 4 k 3.5 f | 14.6 | 14.5 | 14.6 | 685 | Damp. |
|  | 18 | 13.5 | 1 I | c 5 T 4 k 4 f | 15.3 | 15.4 | 15.4 | 690 |  |
|  | 28 | 12.5 | I DDD | b 2 e 3 T 2 f | 23.7 | 23.7 | 23.7 | 700 | g hardly vis. |

8230 S Aquarii Series I-
(1900) $22^{\mathrm{h}} 51^{\mathrm{m}} 45^{\mathrm{a}}\left(+3^{\mathrm{s}} .22\right) ;-20^{\circ} 52^{\prime} .6 \quad\left(+0^{\prime} .32\right)$

Period: $279^{\text {d }} .7$; Variation: $8^{\mathrm{N}}-<12 \frac{1}{2}^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | :---: | :---: |
|  |  |  |  |  |
| d | 2 | $-21^{\circ} 6334$ | 0.0 | 7.9 |
| e | 3 | 6333 | 8.4 | 8.3 |
| a | 4 | 6325 | 10.4 | 8.5 |
| h | 5 | 6341 | 11.1 | 8.8 |
| g | - | 6317 | 16.5 | $(8.9) \mathrm{BD}$. |
| k | 7 | 6342 | 21.0 | 9.4 |
| b | 10 | 6323 | 22.7 | 9.6 |
| m | 11 | 6332 | 25.7 | 9.6 |
| c | 9 | 6336 | 25.7 | 9.5 |
| n | 17 | $-21^{\circ} 6335$ | 26.7 | 10.0 |
|  |  |  |  |  |


| 1800 |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 89 Sept. | 19 | $13^{\mathrm{h}}-15^{\text {n }}$ | I | b 0 S 1 c | 23.7 |  | 23.7 | 11265 |  |
|  | 21 | 15 | I | b 0 S 2 c | 23.2 |  | 23.2 | 267 |  |
|  | 27 | 14.8 | I | a 5 S 3 b ! | 17.6 | 18.1 | 17.9 | 273 |  |
|  | 28 | 13-15 |  | a 5 S 3 b | 17.6 | 18.1 | 17.9 | 274 |  |
| Oct. | 10 | 13 | III DDD | d5S2e1a | 5.7 | 6.0 | 5.9 | 286 |  |
|  | 11 | 13 |  | d 5S 3 a S 2 e | 6.2 6.4 | 6.5 | 6.4 | 287 |  |
|  | 15 | 15.7 | II | d 6 S 3 e | 5.7 | 5.6 | 5.7 | 291 |  |
|  | 17 | 14.7 | II | d 6 S 3 e | 5.7 | 5.6 | 5.7 | 293 |  |
|  | 19 | 15 | III | d 6 S 3 e | 5.7 | 5.6 | 5.7 | 295 |  |
| Nov. | 3 | $13^{h}-15^{\text {a }}$ | II DDD | d 6 S 3 e | 5.7 | 5.6 | 5.7 | 310 |  |
|  | 10 | 14 | III D | d 5S3e | 5.2 | 5.2 | -5.2 | 317 |  |
|  | 14 | 15 |  | d 5 S 4 e | 4.7 | 4.7 | 4.7 | 321 |  |
|  | 28 29 | ${ }_{12}^{15}$ | I D | e2S 1 a | 9.7 11.4 | 9.7 | 9.7 11.4 | 335 | Near horizon, difficult. |
|  | 29 | 12.6 |  | e 15 | 11.4 |  | 11.4 | 336 |  |
|  | 30 | 12.2 |  | a 3 S 7 b | 14.6 | 14.1 | 14.4 | 337 |  |
| Dec. | 1 | 12 |  | a 3 S 7 b | 14.6 | 14.1 | 14.4 | 338 |  |
|  | 11 | 14.3 | I | a 5 S 5 b | 16.6 | 16.6 | 16.6 | 348 | Near horizon |
|  | 12 | 12.7 | I | a 5 S 5 b | 16.6 | 16.6 | 16.6 | 349 |  |
|  |  |  |  | $\mathrm{S}=\mathrm{g}!$ | 16.5 |  |  |  |  |
|  | 21 | 12.7 | 1 | a 4 c 3 S 3 b S 4 e ! | 19.6 21.7 | 19.6 | 20.3 | 358 |  |
|  | 22 | 13 | I | g 3 S 4 b | 19.1 | 19.2 | 19.2 | 359 |  |
|  | 26 | 12 | I | g 5 S 1 b | 21.6 | 21.7 | 21.7 | 363 | Near ${ }^{\text {D }}$ |
| 90 Jan. | 8 | 13 | I | $\mathrm{S}<\mathrm{b}$ |  |  | $\bigcirc 23$ | 376 | Too low, windy. |
| July | 7 | 15.2 | I | $\mathrm{S}<\mathrm{a}$ |  |  | >11 | 556 |  |
|  | 11 | 15.5 | ${ }_{\text {I }}$ | b 2 S 4 c | 23.2 |  | 23.5 | 560 |  |
|  | 14 | 14.3 | II | $\begin{aligned} & \operatorname{losik} \\ & (\mathrm{e} 10 \mathrm{~S} 8 \mathrm{~b}) \end{aligned}$ | $\underset{(16.6)}{18.6}$ | $\begin{gathered} 19.6 \\ (16.4) \end{gathered}$ | 18.2 | 563 | Approximate. |
|  | 16 | 14 | I | $\begin{aligned} & \text { h } 3 \text { S } 5 \mathrm{k} \\ & \text { (S } 10 \mathrm{~b}) \end{aligned}$ | $\left(\begin{array}{l} 15.1 \\ (12.7) \end{array}\right.$ | 14.8 | 14.5 | 565 |  |
|  | 18 | 15.5 | I | h 4 S 3 g | 14.3 | 14.2 | 14.3 | 567 |  |
|  | 20 | 14.6 | D2D | h3S2g | 14.3 | 14.3 | 14.3 | 569 |  |
| Aug. | 3 6 | ${ }_{4} 3^{n}-15^{\text {n }}$ | 22D | e 7 S 5 g | 13.5 | 12.1 | 13.3 | 583 |  |
|  |  |  |  | e 6 S 6 g | 12.5 | 12.5 | 12.5 | $58 ;$ | Difficult. |
|  | 12 | " |  | e 6 S 7 g e 6 S 4 h | 12.0 | 12.1 | 11.2 | 592 |  |
|  | 15 | " |  | e6S4 | 10.8 11.5 | 10.0 | 11.1 | 595 | Seeing very poor. |
|  |  |  |  | e 6 S 4 h | 10.8 | 10.0 |  |  |  |
|  | 23 | 16 |  | e 2 a 6 S 2 h 4 g | 12.8 | 10.9 | 11.9 | 603 |  |
|  | 25 | 17.3 | II D | h 2 S 6 6 | 11.8 | 12.5 | 12.6 | 605 | $\mathrm{h}>\mathrm{S}$ ! |
| Sept. | 3 | 14.2 | I | h2S 1 g ! | 14.3 | 14.7 | 14.5 | 614 |  |
|  | 8 | 17.1 | III | h 4 S 4 k | 16.1 | 16.1 | 17.0 | 619 |  |
|  | 9 |  |  | g 2 S 4 k | 17.8 | 18.0 |  |  |  |
|  | 9 | 14.5 | III | h 6 S 4 k | 17.1 | 17.0 | 17.9 | 620 |  |
|  | 15 | 14.3 | I | g 4S6b, 4 c | 18.6 | 19.0 |  | 626 |  |
|  | 17 | 15.3 | I | g 6 S 2 k | 20.8 | 19.9 | 20.4 | ${ }_{628}^{628}$ |  |
| Oct. | 3 | 13.4 | 1 | b3m3s5n | 25.2 | 26.1 | 25.7 | 644 |  |

## Nova Aurigæ 1892

$$
\begin{gathered}
(1900) 5^{\mathrm{n}} 25^{\mathrm{m}} 34^{\mathrm{s}} \quad\left(+3^{\mathrm{s}} .85\right) ; \quad+30^{\circ} 22^{\prime} .2 \quad\left(+0^{\prime} .05\right) \\
\text { Variation: } 4 \frac{1}{2}^{\mathrm{n}}-<13^{\mathrm{m}} \\
\text { Comparison Stars : }
\end{gathered}
$$

| Obs. | BD . | Steps | BD . | H. | L. |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| a | $+33^{\circ} 1000$ | 0.0 | 5.1 | 5.1 |  |  |
| b | 32922 | 6.0 | 5.5 | 5.3 | \% |  |
| $\chi$ | 321024 | 6.9 | 4.8 | 5.4 | 5.00 |  |
| c | 331013 | 11.1 | 5.9 | 5.5 |  |  |
| d | 30963 | 18.0 | 6.0 | 5.9 | 5.70 |  |
| e | 30898 | 24.7 | 6.2 | 6.2 | 5.86 |  |
| g | $29 \quad 947$ | 29.4 | 6.2 | 6.4 |  |  |
| h | 29899 | 36.2 | 7.0 | 6.6 |  |  |
| k | $29 \quad 911$ | 53.7 | 7.5 | 7.4 |  |  |
| $\alpha$ | $29-923$ | 59.7 | 7.8 | 7.8 |  |  |
| $\beta$ | $29 \quad 921$ | 66.7 | 8.5 | 8.3 |  |  |
| $\gamma$ | $30 \quad 912$ | 68.2 | 8.5 | 8.4 |  |  |
| $\delta$ | 30913 | 75.2 | 8.7 | 8.9 |  |  |
| $\varepsilon$ | $30 \quad 914$ | 85.4 | 9.4 | 9.5 |  |  |
| $\zeta$ | 30920 | 91.7 94.2 | 9.5 | 9.9 10.0 |  | South pr $n$ |
| $\eta$ | +30 924 | 98.9 | 9.5 | 10.3 |  |  |
| $\vartheta$ | ........... | 102.9 | ...... | 10.6 |  | North foll. Nova. |

Notes:
The comparison star $f$, which was used only a few times with the naked eye, has been discarded in the reductions. It consists of two components BD. $+29^{\circ} 953$ and 954 , of magnitude 7.0 and 7.5 respectively, and for this reason its estimates do not agree well among themselves. The space line across the above table separates the naked-eye comparison stars from the telescopic ones. The magnitudes under H were computed by the formulas (see Astr. J. XI, 1892, p. 172):

$$
\begin{aligned}
& \text { Magn. }=6.0+0.045(\text { Steps }-20.7) \text { for opera glass, } \\
& \text { Magn. }=8.6+0.062(\text { Steps }-71.0) \text { for telescope. }
\end{aligned}
$$

The three magnitudes under $L$ are those upon which Lindemann based his definitive light curve of the Nova in the Mélanges Math. et Astr. (Petersburg. Bulletin, t. VII, p. 331).

| Gr. M. T. Sky | Comparisons | I | II | Mean |
| :--- | :--- | :--- | :--- | :--- | :--- |

METHOD BY STEPS:


| $1800+$ |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
|  |  |  |  |  |  |  |  |  |  |

## IV. Observations made by James F. Dawson, S. J., from I889 to 1890.

These observations were made at the Georgetown College Observatory. The instrument employed was a 3 -inch telescope, mounted equatorially but not sheltered under a dome. It had to be carried every evening to a pier in the open air, and adjusted. The variables selected are all southern except Algol, which was observed for practice. The letters (D) and (H) after an observation designate the Rev. J. Daugherty and Hagen respectively who occasionally took part in the observations. The results have been published in the Astronomical Journal.

100
T Ceti Series V
(1900) $0^{\mathrm{h}} 16^{\mathrm{m}} 42^{\mathrm{s}}\left(+3^{\mathrm{s}} .04\right) ;-20^{\circ} 36^{\prime} .7 \quad\left(+0^{\prime} .33\right)$

Period: Irreg.; Variation: $5^{\mathrm{M}}-6 \frac{1}{2}^{\mathrm{M}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | :--- |
|  |  |  |  |  |
| d | 8 | $-19^{\circ} 21$ | 0.0 | 4.8 |
| a | 9 | $21^{\circ} 24$ | 10.0 | 6.4 |
| b | 10 | $19^{\circ} 30$ | 11.0 | 6.6 |
| c | $\ldots . .$. | $-20^{\circ} 48$ | 14.5 | $(7.8) \mathrm{BD}$. |
|  |  |  |  |  |

Notes:
Since this star is "irregularly periodic," the maximum brightness which can be deduced from these observations will be of special importance.

The observations indicate a slight change in the relative brightness of the two comparison stars $a$ and $b$.

| $1800+$ |  | Gr. M. T. | Sky | Comparisons | I | II | Mean | $2400000+$ | Remarks |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 89 Sept. | 22 | 15.5 | II | b 1 T 2 c | 12.3 | 12.2 | 12.3 | 11268 |  |
|  | 27 | 14.5 | I | b0 T 2 c | 11.8 |  | 11.8 | 11273 | *) |
|  | 28 | 13.5 | I | b 2 T 2 c , or | 12.8 | 12.8 | 12.5 | 274 | *) |
| Oct. |  |  |  | b1T2c ${ }_{\text {a }}$ | 12.2 11.0 | 12.2 |  |  | Better. |
|  | 11 | 14.0 | 2 | a 2 T 1 b (H) | 11.0 11.0 | 10.7 | 10.8 | 7 |  |
|  |  |  |  | T 4 c (H) | 10.5 |  |  |  |  |
|  | 15 | 15.5 | I | a 2 T 1 b | 11.0 | 10.7 | 10.6 | 291 |  |
|  |  |  |  | T 5 c | 9.5 |  |  |  |  |
|  | 17 | 15.5 | II | a 1 T1b | 10.5 | 10.5 | 10.5 | 293 |  |
|  | 18 | 16.5 | III | a 1 T 1 b | 10.5 | 10.5 | 10.5 | 294 |  |
|  | 19 | 15 | III | a 1 T1b | 10.5 | 10.5 | 10.5 | 295 |  |
| Nov. | 3 | 13.5 | DDD | d9T1a2b | 9.0 | 9.0 | 8.5 | 310 |  |
|  |  |  |  | d8T2a 1 b (H) | 8.0 | 8.0 |  |  |  |
|  | 10 | 14 | D | d7T3a1b | 7.0 6.0 | 7.0 6.0 | 6.5 | 317 |  |
|  | 14 | 15.5 | II | d7T3a1b | 7.0 | 7.0 | 7.0 | 321 |  |
|  | 15 | 16.0 | I | d 6 T 4 al 1 b | 6.0 | 6.0 | 6.0 | 322 |  |
|  | 16 | 14.0 | III | d6T4a 1 b | 6.0 | 6.0 | 6.0 | 323 |  |
|  | 23 | 15.0 | I | d5T5a0b | 5.0 | 5.0 | 5.0 | 330 |  |
|  | 25 | 15.5 | 1 | d5T5b1a | 5.0 | 5.0 | 5.0 | 332 |  |
|  | 28 | 15.5 | II | d 6 T 5a 1 b | 5.5 | 5.5 | 5.5 | 335 |  |
|  | 30 | 14.0 | $D^{2}$ | d5T5a1b | 5.0 | 5.0 | 5.0 | 337 |  |
| Dec. | 1 2 | 14.5 14.5 | D |  | 4.0 3.0 | 4.0 3.0 | 4.0 3.0 | 338 339 |  |
|  | 4 | 14.5 | D | d 4 T 6 a | 4.0 | 4.0 | 4.0 | 341 |  |
|  | 11 | 14.5 | 1 I | d 3 T 7 a 1 b | 3.0 | 3.0 | 3.0 | 348 |  |
|  | 12 | 12.5 | I | d 3 T 7 a | 3.0 | 3.0 | 3.0 | 349 |  |
|  | 13 | 14.3 | II | d4T6a0b | 4.0 | 4.0 | 4.0 | 350 |  |
|  | 14 |  | II | d 3 T 7 a | 3.0 | 3.0 | 3.0 | 351 |  |
|  | 20 | 14.3 | I | d 4 T 6 a | 4.0 | 4.0 | 4.0 | 357 |  |
|  | 21 | 14.5 | II | d 4 T 6 b | 4.0 | 4.0 | 4.0 | 358 |  |
|  | 22 | 13.5 | II | d 3T7a(?) | 3.0 | 3.0 | 3.0 | 359 |  |
|  | 23 | 14.0 | II | d 4T6a | 4.0 | 4.0 | 4.0 | 360 |  |
|  | 24 | 14.0 | III | d 3 T 7 a 0 b | 3.0 | 3.0 | 3.0 | - 361 |  |
|  | 26 |  |  | d3T7a0b | 3.0 | 3.0 | 3.0 | - 363 |  |
| 90 Jan. | 8 12 | 15.0 | DDD | d 4T6a0b | 4.0 5.5 | 4.0 5.5 | 4.0 5.5 | 376 380 |  |
|  | 13 | 13.0 | I | d 6 T 4 b 1 a | 6.5 | 6.6 | 6.6 | 381 |  |
|  | 16 |  |  | d 6 T4b1a | 6.5 | 6.6 | 6.6 | 384 |  |
|  | 17 |  | III | d6T4a0b | 6.0 | 6.0 | 6.0 | 385 |  |
|  | 21 |  | II | d 7 T 4 b 1 a | 7.0 | 7.0 | 7.0 | 389 |  |
|  | 27 28 |  | ${ }_{D}$ |  | 6.5 7.0 | 6.6 7.0 | 6.6 7.0 | 395 396 |  |
|  | 28 |  | ) |  |  |  |  |  |  |

*) The original has $d$ instead of $c$ on both nights. Considering the observations before and after, and the magnitude of $d$, the correction becomes almost certain.
$(1900) 3^{\mathrm{h}} 1^{\mathrm{m}} 40^{\mathrm{s}}\left(+3^{\mathrm{s}} .89\right) ;+40^{\circ} 34^{\prime} .2 \quad\left(+0^{\prime} .23\right)$
Period: $2^{\mathrm{a}} 20^{\mathrm{h}} 48^{\mathrm{m}} .9$; Variation: $2 \frac{1}{2}^{\mathrm{N}}-3 \frac{1}{2}^{\mathrm{N}}$.
Comparison Stars:

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
| $\gamma$ Persei | 19 | $+52^{\circ} 654$ | 0.0 | 3.0 |
| o " | 26 | $47^{\circ} 876$ | 3.0 | 3.1 |
| $\kappa$ | 23 | $44^{\circ} 631$ | 8.0 | 4.1 |
| $r$ Androm. | 11 | $41^{\circ} 395$ | 0.0 | 2.3 |
| $\zeta$ Persei | 30 | $31^{\circ} 666$ | 9.0 | 2.9 |
| o " | 26 | $47^{\circ} 876$ | 10.4 | 3.1 |
| $\alpha$ Triang. | 37 | $28^{\circ} 312$ | 12.4 | 3.6 |
| $\rho$ Persei | 20 | $38^{\circ} 630$ | 16.4 | Var. |
| $\kappa$ " | 23 | $+44^{\circ} 631$ | 22.2 | 4.1 |

## Notes:

The two scales of comparison stars refer to the two evenings on which the observations were made. The latter were intended only as an exercise in estimating differences in magnitude, but may be useful as a confirmation of simultaneous observations of the same minima made elsewhere.

(1900) $4^{\mathrm{n}} 55^{\mathrm{m}} 3^{\mathrm{s}}\left(+2^{\mathrm{s}} .73\right) ;-14^{\circ} 57^{\prime} .4 \quad\left(+0^{\prime} .09\right)$

Period: $436^{\mathrm{d}} .1$; Variation: $6 \frac{1}{2}^{\mathrm{N}}-8 \frac{1}{2}^{\mathrm{N}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | ---: | ---: | ---: | ---: |
|  |  |  |  |  |
| e | 1 | $-14^{\circ} 1003$ | 0.0 | 6.7 |
| a | 2 | $15^{\circ} 910$ | 6.0 | 7.6 |
| b | 9 | $14^{\circ} 1005$ | 13.0 | 8.6 |
| c | 12 | $15^{\circ} 912$ | 16.0 | 8.9 |
| d | 20 | $-14^{\circ} 1009$ | 21.0 | 9.4 |
|  |  |  |  |  |

Notes:
Observations of this star are difficult both on account of its redness and of the length of its period. The following observations will give the ascending branch of the light curve, and may serve as a supplement to observations made elsewhere.


2610 R Canis Maioris Series V.
(1900) $7^{\mathrm{h}} 14^{\mathrm{m}} 56^{\mathrm{s}}\left(+2^{\mathrm{s}} .70\right) ;-16^{\circ} 12^{\prime} .4$ ( $-0^{\prime} .11$ )

Period: $1^{\mathrm{d}} 3^{\mathrm{h}} 15^{\mathrm{m}} .8$; Variation: $6^{\mathrm{M}}-6 \frac{1}{2}^{\mathrm{m}}$.
Comparison Stars.

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| a | 28 | $-15^{\circ} 1734$ | $0.0,0.0,0.0$ <br> b | - |
| $-15^{\circ} 1732$ | $6.8,7.4,9.7$ | 5.9 HP <br> $(6.8) \mathrm{BD}$. |  |  |

Notes:
The three scales in the column "Steps" refer to the three Minima. A mean scale would not represent the observations as well as these deduced from the three sets separately. A glance at Chart VII of Series V. of the Atlas shows that better comparison stars than $b$ could have been ehosen.


## 2676 U Monocerotis Series IV.

(1900) $7^{\mathrm{h}} 26^{\mathrm{m}} 1^{\mathrm{s}}\left(+2^{\mathrm{s}} .86\right) ;-9^{\circ} 34^{\prime} .0 \quad\left(-0^{\prime} .12\right)$

Period: $46^{\mathrm{d}} .10$; Variation : $6 \frac{1}{2}^{\mathrm{M}}-7^{\mathrm{M}}$.
Comparison Stars :

| Obs. | ASV. | BD. | Steps | Magn. |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| a | 1 | $-10^{\circ} 2067$ | 0.0 | 5.8 |
| b | 3 | $9^{\circ} 2086$ | 5.2 | 6.6 |
| c | 4 | $9^{\circ} 2069$ | 7.2 | 6.8 |
| e | - | $9^{\circ} 2043$ | 10.2 | $(7.0) \mathrm{BD}$. |
| f | 8 | -92084 | 16.2 | 7.8 |


*) Perhaps: a 5 U 1 b ?

## CONTENTS.

List of the 52 Variable Stars Observed.

| Star |  | Section | Page |  | Star | Section | Page |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 100 | T Ceti | IV | 136 | 4847 | S Virginis | III | 125 |
| 320 | U Cephei | I | 6 | 5157 | S Bootis | I | 43 |
| 782 | R Arietis | II | 9 | 5374 | ¢ Libræ | II | 87 |
| 806 | - Ceti | III | 114 | 5484 | U Coronæ | I | 46 |
| 814 | S Persei | I, III | 13, 115 | 5501 | S Serpentis | I | 49 |
| 845 | R Ceti | III | 116 | 5504 | S Coronæ | I | 51 |
| 893 | U Ceti | III | 118 | 5770 | R Herculis | I | 54 |
| 976 | T A Arietis | 1 | 17 | 5950 | W Herculis | I | 56 |
| 1090 | $\beta$ Persei | I, IV | 21, 138 | 6044 | S Herculis | I | 58 |
| 1222 | R Persei | I | 23 | 6132 | R Ophinchi | III | 127 |
| 1411 | 2 Tauri | II | 80 | 6181 | a Herculis | II | 89 |
| 1771 | R Leporis | IV | 140 | 6202 | u Herculis | II | 92 |
| 1855 | R Aurigæ | I | 26 | 6512 | T Herculis | I | 61 |
| 2098 | ${ }^{\circ}$ Orionis | II | 82 | 6758 | $\beta$ Lyræ | II | 94 |
| 2100 | U Orionis | I | 29 | 6905 | R Sagittarii | III* | 128 |
| 2509 | \% Geminorum | II - | 84 | 6921 | S Sagittarii | III | 129 |
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| 2610 | R Canis Mai. | IV | 142 | 7106 | S Vulpeculæ | I | 66 |
| 2676 | U Monocerotis | IV | 144 | 7120 | $\chi$ Cygni | I | 70 |
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| 3109 | S Cancri | I | 35 | 7257 | R Sagittæ | I | 72 |
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| 3825 | R Ursæ Mai. | I | 40 | 7468 | T Aquarii | III | 130 |
| 4407 | R Corvi | III | 120 | 7803 | $\mu$ Cephei | II | 99 |
| 4805 | W Virginis | III | - 122 | 8073 | ${ }^{\text {o }}$ Cephei | II | 103 |
| 4816 | V Virginis | III | 124 | 8230 | S Aquarii | III | 131 |

List of 5 Stars not Strictly variable.

| Star | Section | Page | Star | Section | Page |
| :--- | :---: | :---: | :--- | :---: | :---: |
| Nova Andromedæ <br> Nova Aurigæ <br> $\delta$ Serpentis | I | 78 | (W) Bootis | II | 107 |
| 133 | P Cygni | II | 1 II |  |  |

Additions and Corrections.
Note to page 4: In case of a few variables occurring in Series IV the designations and magnitudes of the comparison stars have been inserted from the MS. of that Series, which is now nearly ready for print.

| Page | $1800+$ |  | Columns | Corrections. |
| :---: | :---: | :---: | :---: | :---: |
| 10 | 83 | Mar. 31 | Mean | insert: $>22$ |
| 53 |  | " 17 | ، | read: 4.7 |
| 113 | 87 | Jan. 28 | I " | insert: 2.0 |
| 121 |  | May 3 | I ، | cancel second 16.5 |
| 141 |  | Mar. 23 | I. " | insert: 0.0 |



$\qquad$


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$$
\begin{aligned}
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& \text { (11) (11) (11) (11) (11) } \\
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& \text { (11) (11) (112) (112) } \\
& \text { (11) }(11)(11)=212) \\
& \text { (112) (11) (11) (11) (11) } \\
& \text { (11) (11) (11) ( (12) (11) }
\end{aligned}
$$



